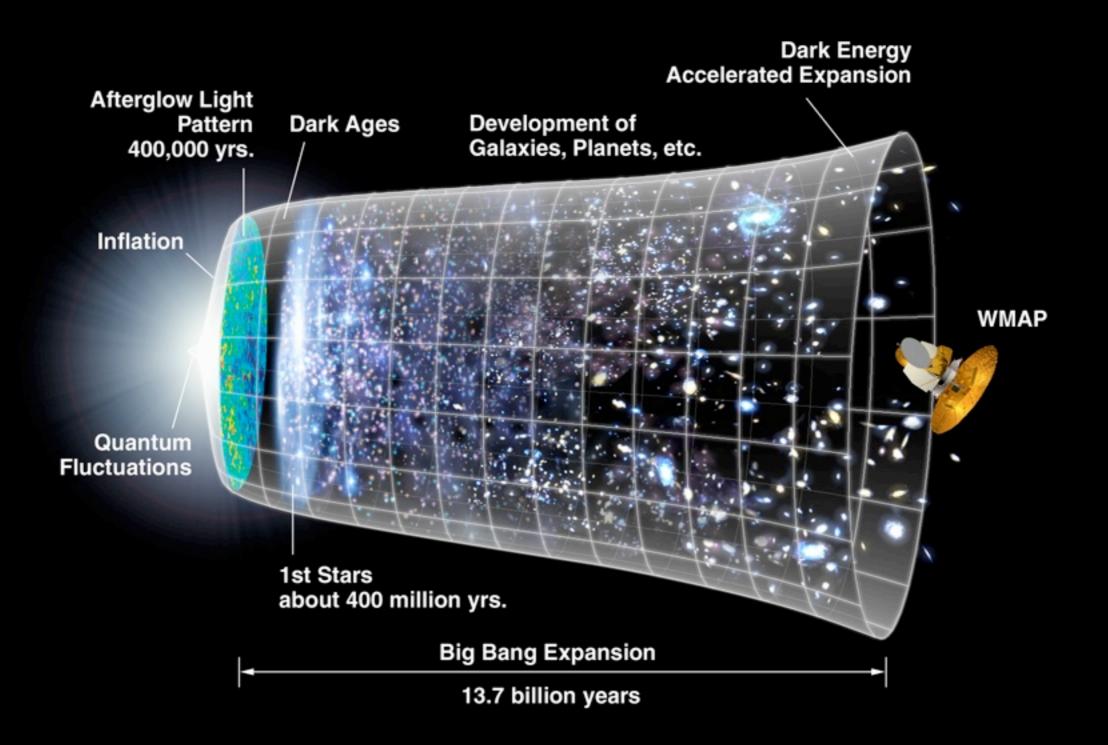
#### Observational Evidence for Dark Energy: Summary and Assorted Thoughts

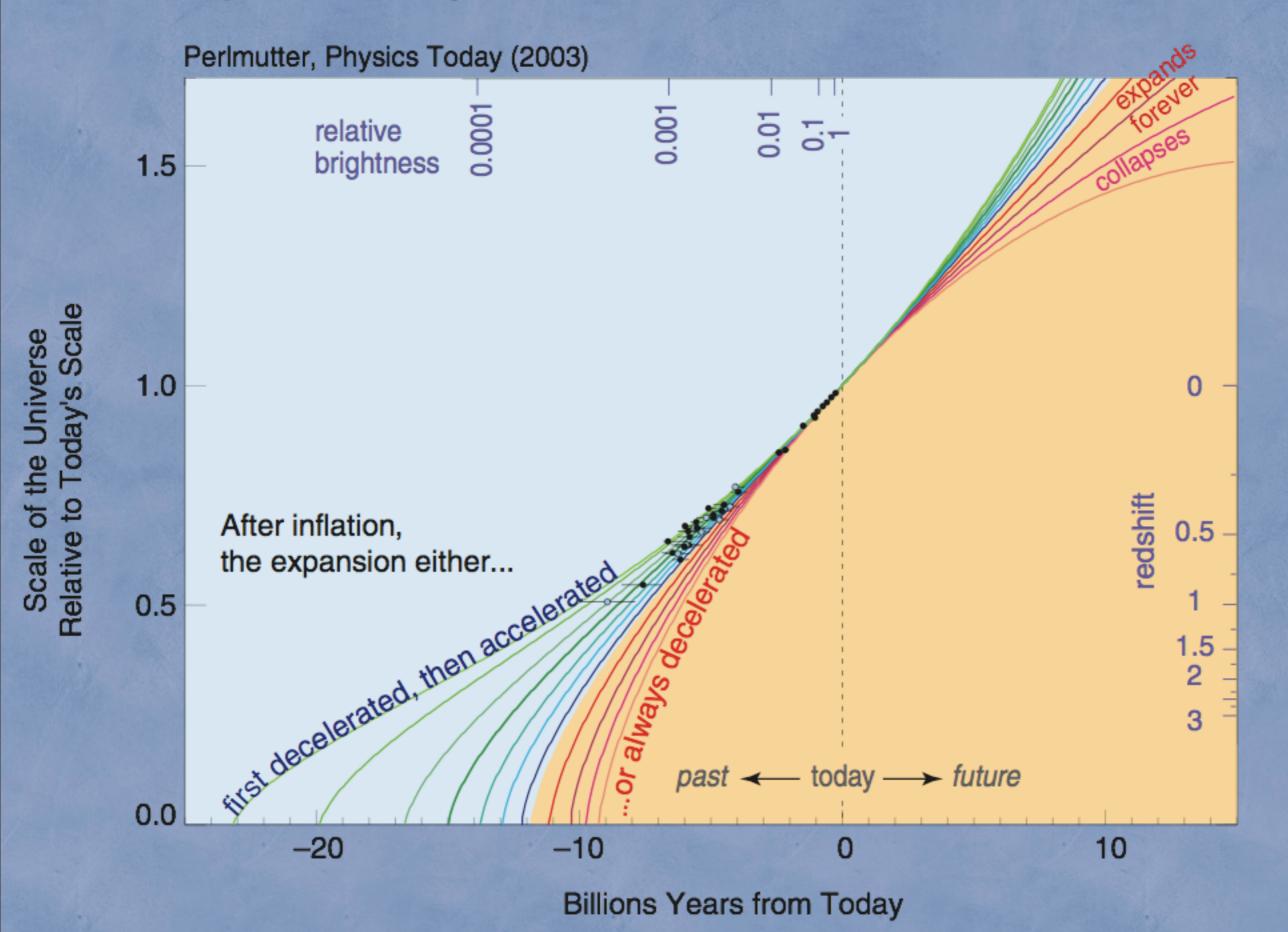
Michael Wood-Vasey Harvard University

IPM Cosmology School & Workshop

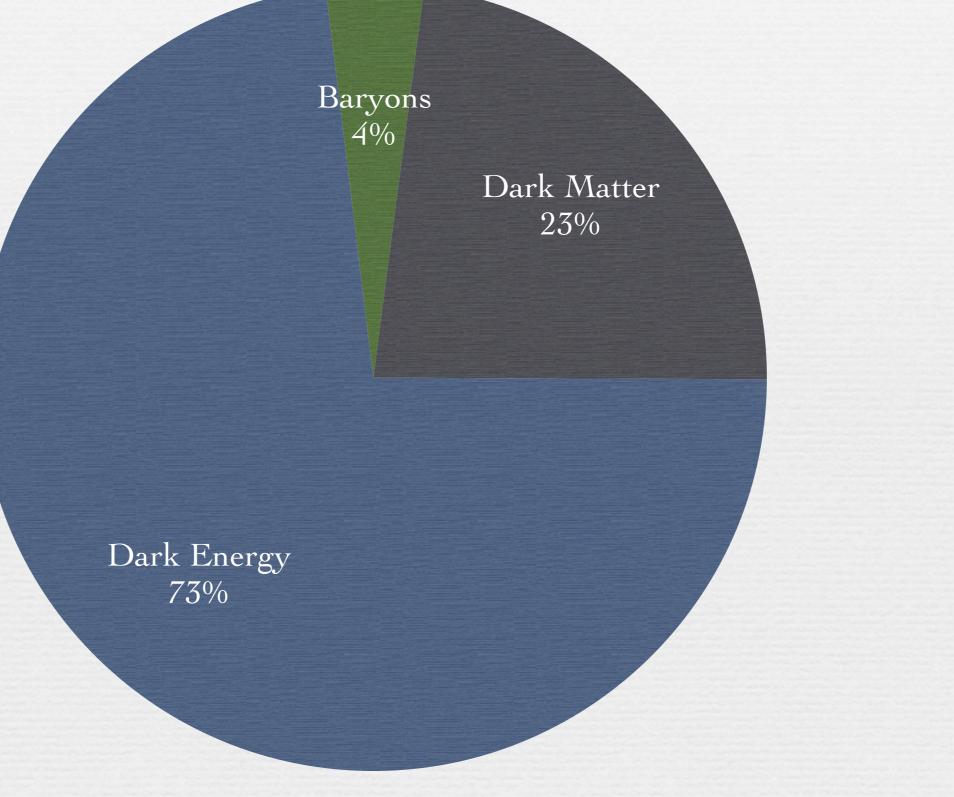
2007 June 8



#### **Expansion History of the Universe**

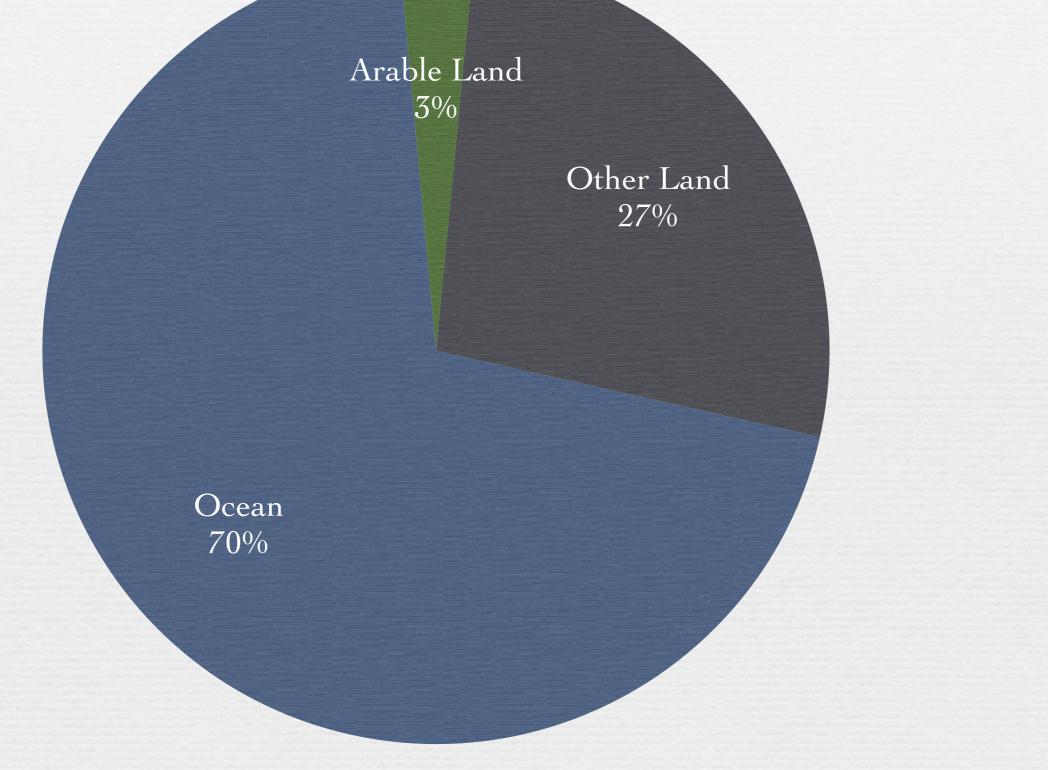


### Composition of Universe



c.f. Astier06, Spergel06, Eisenstein05, Perlmutter99, Riess98

#### Composition of Earth's Surface



Food and Agriculture Organization of the United Nations

## What Is Dark Energy?

- Repulsive component
- ✤ Negative pressure
- Everywhere
- No observed structure
- Particle Physics?
- Cosmological Constant?
- Modification of GR?

#### The Basic Question:

Is a cosmological constant model consistent with our observations of the Universe?

#### The Basic Question:

## $I_{S} W = -1?$

#### How should we measure w?

Dynamic Methods to constrain theorists

- Baryon Acoustic Oscillations
- Galaxy Clusters
- SNIa Luminosity Distances
- ✤ Weak Lensing

Basic Framework to test theorists

CMB: curvature, GW -- test inflation

#### Dark Matter

- Have several decent theories
- Could find in next 10 years
- ✤ New particle
  - ✤ LHC
  - recoil experiments
- Astrophysics
  - Self-annihilation signal (e.g., GLAST)
  - baryon filling of DM structure
    - observations from sensitive large surveys
    - improved simulations

## Dark Energy

- Have no decent theories
- Currently appears pervasive
- Maybe has structure?
- Variable with cosmic time
- Observations throughout cosmic time
  - Prioritization tricky w/o better theory
- \$100 million \$1 billion to make progress

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may be related

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Potential donors please contact wmwood-vasey@cfa.harvard.edu

may be related

# All I want is a well-motivated theory!

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#### Is that too much to ask?

Is this really the path towards understanding Dark Energy?

Courtesy of David Weinberg

