Beyond the Higgs new ideas on electroweak symmetry breaking

Higgs mechanism. The Higgs as a UV moderator of EW interactions. Needs for New Physics beyond the Higgs. Dynamics of EW symmetry breaking

Review of possible scenarios :Gauge-Higgs Unification, Little Higgs, Composite Higgs, (5D) Higgsless models.

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A decade of experimental successes

top discovery

Solar and atmospherical neutrino oscillations

O direct CP violation in the K system (ds) (K-long decaying into 2 pions)

 \bigcirc CP violation in the B system (db)

Solution evidence of an accelerated phase in the expansion of the Universe

The measure of the dark energy/dark matter composition of the Universe

These results have strengthened the SM as a successful description of Nature ... but







... these experimental results also concluded that there is a physics beyond the Standard Model.

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Bounds on (Dangerous) New Physics

Heavy Particles \Rightarrow new interactions for SM particles

broken symmetry	operators	$scale\Lambda$
B,L	$(QQQL)/\Lambda^2$	$10^{13} { m TeV}$
flavor $(1,2^{nd} \text{ family}), CP$	$(ar{d}sar{d}s)/\Lambda^2$	$1000 { m TeV}$
flavor $(2,3^{rd} \text{ family})$	$m_b(\bar{s}\sigma_{\mu\nu}F^{\mu\nu}b)/\Lambda^2$	$50 { m TeV}$

At colliders, it will be difficult to find direct evidence of new physics in these sectors...

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New Physics in the EW sector

$\left((h^{\dagger} \sigma^{a} h) W^{a}_{\mu\nu} B^{\mu\nu} \right) / \Lambda^{2} \qquad |h^{\dagger} D_{\mu} h|^{2} / \Lambda^{2} \qquad \left(h^{\dagger} h \right)^{3} / \Lambda^{2}$

$\Lambda \sim {\rm few}~{\rm TeV}~{\rm only}$

high potential for direct detection at LHC, ILC !!!

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EW "unification" and EWSB



Above ~ 100 GeV, electromagnetic and weak interactions are unified

Below 100 GeV, γ and Z behave differently

 $m_{\gamma} < 6 \times 10^{-17} \text{ eV}$ $m_{W^{\pm}} = 80.425 \pm .038 \text{ GeV}$ $m_{Z^0} = 91.1876 \pm .0021 \text{ GeV}$

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Higgs Mechanism



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Beyond the Higgs

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Custodial Symmetry

Rho parameter

$$\begin{pmatrix} \rho \equiv \frac{M_W^2}{M_Z^2 \cos^2 \theta_w} = \frac{\frac{1}{4}g^2 v^2}{\frac{1}{4}(g^2 + g'^2)v^2 \frac{g^2}{g^2 + g'^2}} = 1 \end{pmatrix}$$
Consequence of an approximate global symmetry of the Higgs sector
$$= \begin{pmatrix} H^+ \\ H^0 \end{pmatrix} \text{ Higgs doublet = 4 real scalar fields}$$

$$T(H) = \lambda \left(H^{\dagger}H - \frac{v^2}{2}\right)^2 \text{ is invariant under the rotation of the four real components}}$$

$$SU(2)_R \qquad \Phi^{\dagger}\Phi = H^{\dagger}H \begin{pmatrix} 1 \\ 1 \end{pmatrix}$$

$$SU(2)_R \qquad \Phi^{\dagger}\Phi = H^{\dagger}H \begin{pmatrix} 1 \\ 1 \end{pmatrix}$$

$$V(H) = \frac{\lambda}{4} \left(\operatorname{tr}\Phi^{\dagger}\Phi - v^2\right)^2$$

$$explicitly invariant under SU(2)_L \times SU(2)_R$$

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SU

0

H

Custodial Symmetry

Higgs vev $\begin{array}{l} \langle H \rangle = \begin{pmatrix} 0 \\ \frac{v}{\sqrt{2}} \end{pmatrix} \langle \Phi \rangle = \frac{v}{\sqrt{2}} \begin{pmatrix} 1 \\ 1 \end{pmatrix} \\ \begin{pmatrix} W_{\mu}^{1}, W_{\mu}^{2}, W_{\mu}^{3} \end{pmatrix} \text{ transforms as a triplet} \\ \\ (Z_{\mu} \gamma_{\mu}) \begin{pmatrix} M_{Z}^{2} & 0 \\ 0 & 0 \end{pmatrix} \begin{pmatrix} Z^{\mu} \\ \gamma^{\mu} \end{pmatrix} = (W_{\mu}^{3} B_{\mu}) \begin{pmatrix} c^{2} M_{Z}^{2} & -cs M_{Z}^{2} \\ -cs M_{Z}^{2} & s^{2} M_{Z}^{2} \end{pmatrix} \begin{pmatrix} W^{3 \mu} \\ B^{\mu} \end{pmatrix} \\ \\ \\ \text{The } SU(2)_{V} \text{ symmetry imposes the same mass term for all } W^{i} \text{ thus } c^{2} M_{Z}^{2} = M_{W}^{2} \\ \\ \rho = 1 \end{array}$

More generally:

$$\rho = \frac{\sum_{i} (T_i (T_i + 1) - T_{3i}^2) v_i^2}{2T_{3i}^2 v_i^2}$$

Need doublet (T_i=1/2) with hypercharge 1/2 (T_{3i}=1/2) to get ρ = 1

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The hypercharge gauge coupling and the Yukawa couplings break the custodial SU(2)_{V.} The radiative corrections will generate a (small) deviation to $\rho = 1$ at one loop.

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Counting the degrees of freedom $SU(2)_l \times U(1)_y \rightarrow U(1)_{em}$

3 broken gauge directions = 3 eaten Goldstone bosons

3 eaten Goldstone bosons which become the longitudinal polarizations of the massive gauge bosons

 $H = \left(\begin{array}{c} h^+ \\ h^0 \end{array}\right)$

One physical degree of freedom the Higgs boson

$$egin{aligned} H &= e^{i\pi^a T^a} \left(egin{aligned} 0 \ rac{v+h}{\sqrt{2}} \end{array}
ight) & \mbox{In} \ V(H) &= \lambda \left(H^\dagger H - rac{v^2}{2}
ight)^2 \end{aligned}$$

In the unitary gauge, the π^a are non-physical.

$$= \lambda v^2 h^2 + \lambda v h^3 + \frac{1}{4}\lambda h^4$$

mass term

Higgs doublet = 4 real scalar fields

self-couplings

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Higgs as a UN moderator

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Why do we need a Higgs?

The W and Z masses are inconsistent with the known particle content! Need more particles to soften the UV behavior of massive gauge bosons.

(in the R- ξ gauge, the time-like polarization ($\epsilon^\mu\epsilon_\mu=1,\;k^\mu\epsilon_\mu=M$) is arbitrarily massive and decouple)

Bad UV behavior for the scattering of the longitudinal polarizations



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Why do we need a Higgs?

Bad UV behavior for the scattering of the longitudinal polarizations



 $\mathcal{A} = \epsilon_l^{\mu}(k)\epsilon_l^{\nu}(l) ig^2(2\eta_{\mu\rho}\eta_{\nu\sigma} - \eta_{\mu\nu}\eta_{\rho\sigma} - \eta_{\mu\sigma}\eta_{\nu\rho})\epsilon_l^{\rho}(p)\epsilon_l^{\sigma}(q)$



violations of perturbative unitarity around E ~ M

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Why do we need a Higgs?



 H^0

 W^{-}

 W^+

 W^+ W^+ W^+

 $\mathcal{A} = g^2 \left(\frac{E}{M_W}\right)$



 $\mathcal{A} = g^2 \left(\frac{M_H}{2M_W}\right)^2$

Lewellyn Smith '73 Dicus, Mathur '73 Cornwall, Levin, Tiktopoulos '73

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The Higgs boson unitarize the W scattering (if its mass is below ~700 GeV)

 W^+

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+

 W^+

Unitarity Bound

Lee, Quigg, Thacker '77 Chanowitz, Gaillard '85

 $\frac{d\sigma}{d\Omega} = \frac{1}{64\pi^2 s} \left| \mathcal{A} \right|^2$ Partial wave amplitude decomposition: $a_l = \frac{1}{32\pi} \int_{-1}^{1} d(\cos\theta) P_l(\cos\theta) \mathcal{A}$ $\mathcal{A} = 16\pi \sum (2l+1)P_l(\cos\theta)a_l$ $\sigma = \frac{16}{s} \sum_{l=0}^{\infty} (2l+1)|a_l|^2$ $P_0(x) = 1, P_1(x) = x, P_2(x) = 3x^2/2 - 1/2...$ **Optical theorem:** $\left|\operatorname{Re}\left(a_{l}\right)\right| \leq 1/2$ $a_0 = \frac{g^2 E^2}{16\pi M_W^2}$ $E \le 620 \text{ GeV}$ SM witout a Higgs Stronger Bound $a_0 = \frac{g^2 M_H^2}{64\pi M_W^2}$ $M_H \le 1.2 \text{ TeV}$ $2W^+W^- + ZZ$ scattering SM with a Higgs $M_H \leq 780 \text{ GeV}$ Christophe Grojean Tohran, May '08

Physics Beyond the Higgs?

Is the Standard Model with a Higgs a UV finite theory? i.e. valid to arbitrarily high energies

Of course, the SM will fail around the Planck scale but the real question is

Is there any reason to think there is new physics between the weak scale and the Planck scale?

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Quantum corrections of the Higgs potential







Quantum Behavior of the Higgs⁴ Coupling (I)

$$V(h) = -\frac{1}{2}\mu^{2}h^{2} + \frac{1}{4}\lambda h^{4}$$

 $|vev:v^2=\mu^2/\lambda|$

mass : $m_H^2 = 2\lambda v^2$

 $16\pi^2 \frac{d\lambda}{d\ln Q} = 24\lambda^2 - (3g'^2 + 9g^2 - 12y_t^2)\lambda + \frac{3}{8}g'^4 + \frac{3}{4}g'^2g^2 + \frac{9}{8}g^4 - 6y_t^4 + \frac{1}{8}g'^4 + \frac{1}{8}g'^2g^2 + \frac{9}{8}g^4 - \frac{1}{8}g'^4 + \frac{1}{8}g'^2g^2 + \frac{1}{8}g'^4 + \frac{1}{8}g'^2g^2 + \frac{1}{8}g'^4 + \frac{1}{8}g'^2g^2 + \frac{1}{8}g'^4 + \frac{1}{8}g'^4g^4 + \frac{1}{8}g'^2g^2 + \frac{1}{8}g'^4g^4 + \frac{1}{8}g'^2g^2 + \frac{1}{8}g'^4g^4 + \frac{1}{8}g'^2g^2 + \frac{1}{8}g'^4g^4 + \frac{1}{8}g'^2g^2 + \frac{1}{8}g'^4g^4 + \frac{1}{8}g'^4g^4 + \frac{1}{8}g'^2g^2 + \frac{1}{8}g'^4g^4 +$

Large mass (λ dominated RGE)

$$16\pi^2 \frac{d\lambda}{d\ln Q} = 24\lambda^2$$

 λ increases with Q: IR-free coupling

$$\lambda(Q) = \frac{m_H^2}{2v^2 - \frac{3}{2\pi^2}m_H^2\ln(Q/v)}$$





Quantum Behavior of the Higgs⁴ Coupling (I) $V(h) = -\frac{1}{2}\mu^2 h^2 + \frac{1}{4}\lambda h^4$

 $16\pi^2 \frac{d\lambda}{d\ln Q} = 24\lambda^2 - (3g'^2 + 9g^2 - 12y_t^2)\lambda + \frac{3}{8}g'^4 + \frac{3}{4}g'^2g^2 + \frac{9}{8}g^4 - 6y_t^4 + \frac{4}{3}g'^2g^2 + \frac{9}{8}g^4 - \frac{10}{3}g'^4 + \frac{10}{3}g'^2g^2 +$

Large mass (λ dominated RGE)

$$\lambda(Q) = \frac{m_H^2}{2v^2 - \frac{3}{2\pi^2}m_H^2 \ln(Q/v)}$$



Landau pole

 $\Lambda < v \, e^{4\pi^2 v^2 / 3m_H^2}$

New physics should appear before that point to restore stability

for $m_{\rm H}$ fixed, upper bound on Λ for Λ fixed, upper bound on $m_{\rm H}$

No microscopic description: for $\Lambda o \infty$, trivial theory (λ =0)

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Beyond the Higgs

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Quantum Behavior of the Higgs⁴ Coupling (II) $16\pi^2 \frac{d\lambda}{d\ln Q} = 24\lambda^2 - (3g'^2 + 9g^2 - 12y_t^2)\lambda + \frac{3}{8}g'^4 + \frac{3}{4}g'^2g^2 + \frac{9}{8}g^4 - 6y_t^4 + \frac{1}{8}g'^4 + \frac{1}{8}g'^2g^2 + \frac{9}{8}g^4 - 6y_t^4 + \frac{1}{8}g'^4 + \frac{1}{8}g'^2g^2 + \frac{9}{8}g^4 - \frac{1}{8}g'^4 + \frac{1}$

Small mass (yt dominated RGE)

$$16\pi^2 \frac{d\lambda}{d\ln Q} = -6y_t^4$$

 λ decreases with Q.

$$\left(16\pi^2 \frac{dy_t}{d\ln Q} = \right)$$

 $\frac{9}{2}y_t^3 + rac{1}{2}$ Higher loops $y^2(Q) = rac{y^2(Q_0)}{1 - rac{9}{16\pi^2}y^2(Q_0)\lnrac{Q}{Q_0}}$

$$\lambda(Q) = \lambda_0 - \frac{\frac{3}{8\pi^2} y_0^4 \ln \frac{Q}{Q_0}}{1 - \frac{9}{16\pi^2} y_0^2 \ln \frac{Q}{Q_0}}$$

Christophe Grojean $2\pi^2 m_{II}^2 / 3u_{I}^4 v^2$



Quantum Behavior of the Higgs⁴ Coupling (II) $16\pi^{2} \frac{d\lambda}{d \ln Q} = 24\lambda^{2} - (3g'^{2} + 9g^{2} - 12y_{t}^{2})\lambda + \frac{3}{8}g'^{4} + \frac{3}{4}g'^{2}g^{2} + \frac{9}{8}g^{4} - 6y_{t}^{4} + \frac{\text{Higher loops}}{\text{Small Yukawa}}$

Small mass (yt dominated RGE)

$$\lambda(Q) = \lambda_0 - \frac{\frac{3}{8\pi^2} y_0^4 \ln \frac{Q}{Q_0}}{1 - \frac{9}{16\pi^2} y_0^2 \ln \frac{Q}{Q_0}}$$

 $\lambda < 0 \qquad \Rightarrow$ potential unbounded from below

$$\lambda(Q)=0$$
 for $\lambda_0pprox rac{3}{8\pi^2}y_0^4\lnrac{Q}{Q}$

New physics should appear before that point to restore stability

$$\Lambda \le v \, e^{4\pi^2 m_H^2 / 3y_t^4 v^2}$$

for Λ fixed, lower bound on m_{H}



v

 λ

 $\frac{m_H^2}{2v^2}$

0

Beyond the Higgs

Q

 $v e^{4\pi^2 m_H^2/3y_t^4 v^2}$





the SM is not UV complete it is an effective theory of a more comprehensive theory the cutoff of the SM can be rather low

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Solution to the Higgs⁴ Coupling Instabilities

find a symmetry such that



the Higgs quartic will inherit the good UV asymptotically free behavior of the gauge coupling

Examples of such symmetry:

supersymmetry



gauge-Higgs unification: the Higgs is identified as a component of the gauge field along some extra-dimensions.

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Quantum Instability of the Higgs Mass

so far we looked only at the RG evolution of the Higgs quartic coupling (dimensionless parameter). The Higgs mass has a totally different behavior: it is higly dependent on the UV physics, which leads to the so called hierarchy and little hierarchy problems.



A low-mass Higgs boson is imperiled by quantum corrections. The hierarchy problem is a technical problem in scalar theories

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at two loop



Veltman's throat

 $m_H^2 = 4m_t^2 - 2m_W^2 - m_Z^2 \approx (320 \text{ GeV})^2$

(one-loop)

The throat actually closes up at higher loops

shouldn't be taken seriously: in absence of any symmetry, there is no reason to use the same cutof for all particles

Kolda, Murayama '00

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Symmetries for a natural EW/SB







How to Stabilize the Higgs Potential

Goldstone's Theorem

massless scalar spontaneously broken global symmetry ... but the Higgs has sizable non-derivative couplings The spin trick 2s+1 polarization states a particle of spin s: ...with the only exception of a particle moving at the speed of light ... fewer polarization states Spin 1 Gauge invariance \longrightarrow no longitudinal polarization **m=0** Spin 1/2 Chiral symmetry \longrightarrow only one helicity ... but the Higgs is a spin 0 particle

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Symmetries to Stabilize a Scalar Potential

Supersymmetry

fermion ~ boson

 $A_{\mu} \sim A_5$

Higher Dimensional Lorentz invariance

4D spin 1

4D spin 0

These symmetries cannot be exact symmetry of the Nature. They have to be broken. We want to look for a soft breaking in order to preserve the stabilization of the weak scale.

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Other symmetries?

[Grinstein, O'Connell, Wise '07]

Ghost symmetry

SM particle ~ ghost

It was known since Pauli-Villars that ghosts can soften the UV behavior of the propagators. But they are unstable per se.

Lee-Wick in the 60's proposed a trick to stabilize the ghosts (at the price of a violation of causality at the microscopic scale).

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EW interactions need a UV moderator to unitarize WW scattering amplitude

the SM with a light fundamental scalar cannot be (naturally) extended to very high energies

new particles/symmetries are expected to populate the TeV scale to trigger the breaking of the EW symmetry

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