

Resonant leptogenesis

theoretical advances and experimental prospects

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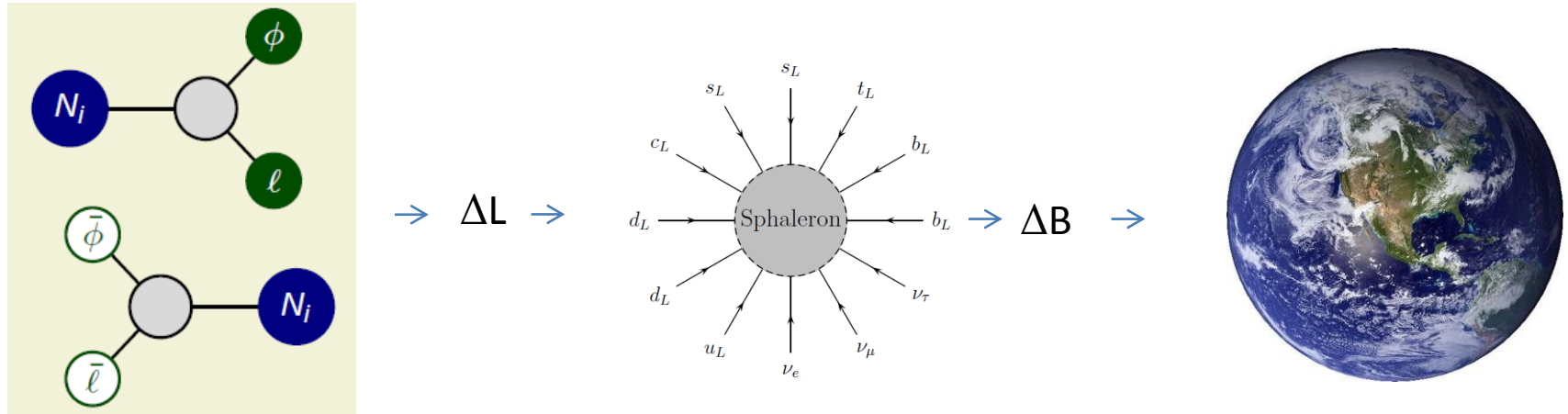
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Outline

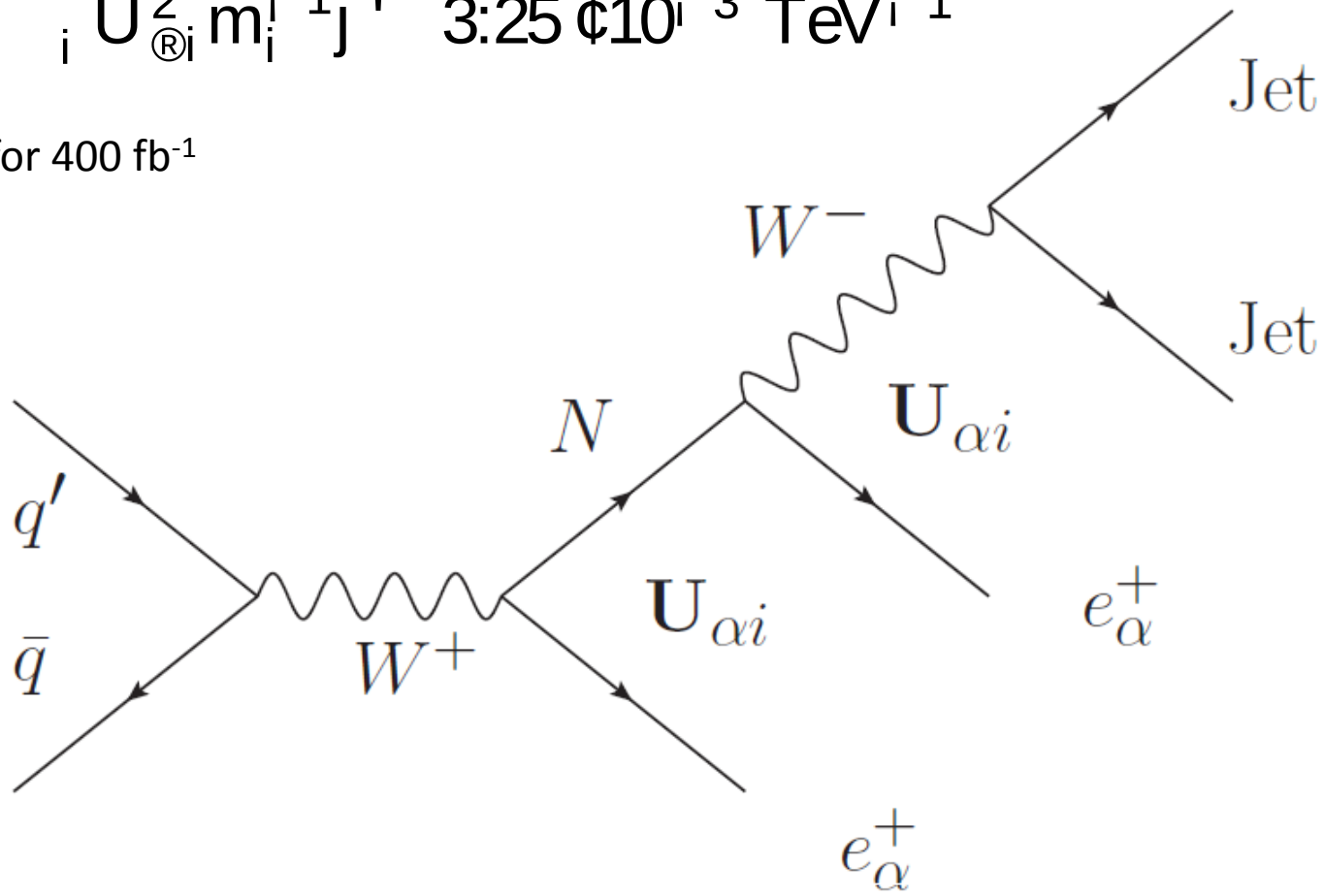
- Leptogenesis via decays
- Collider searches
- Resonant leptogenesis
- Kadanoff-Baym approach
- Leptogenesis via oscillations
- Experimental search
- Summary

Leptogenesis via decays

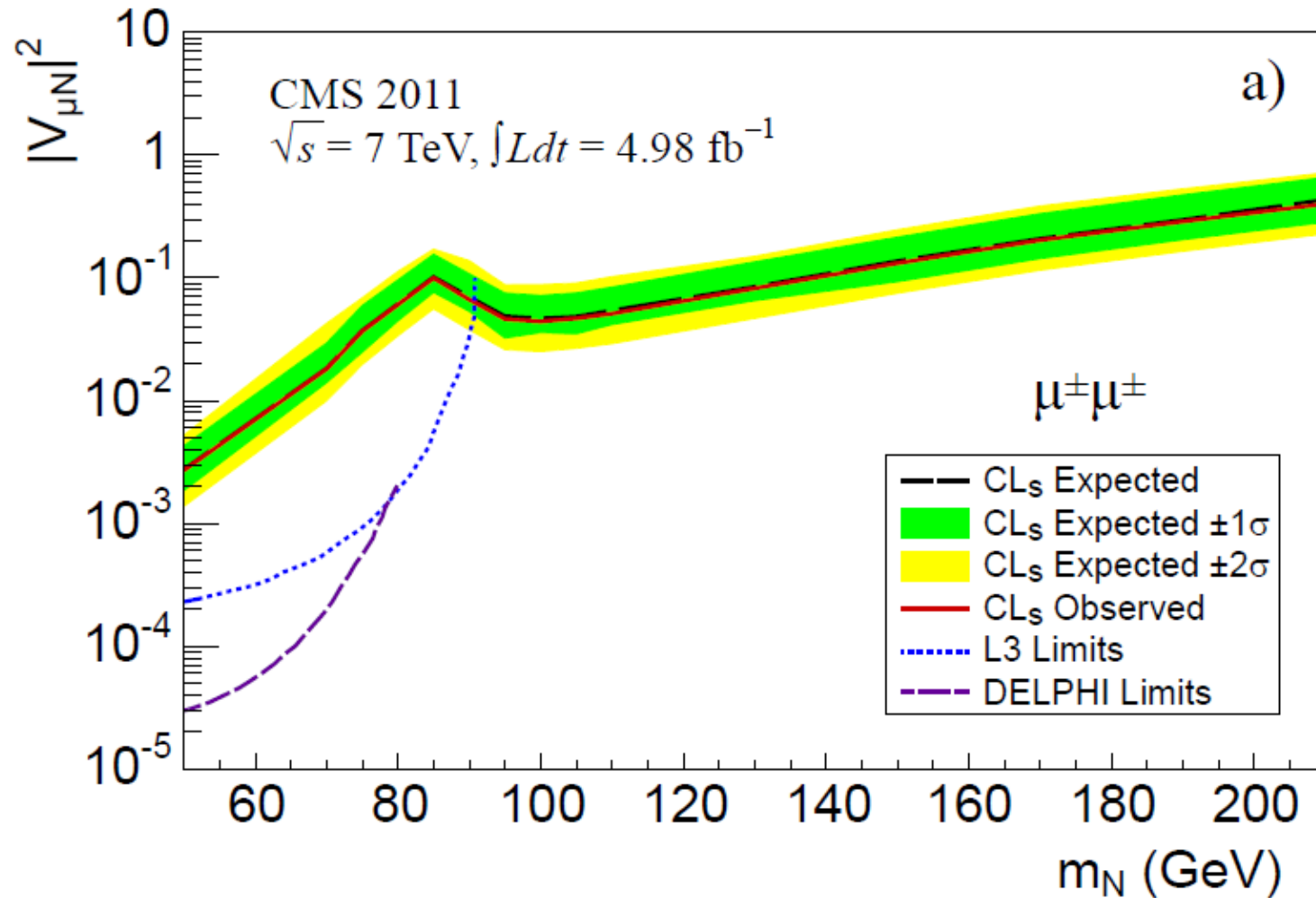


Collider signature

$$j^{\Gamma_i} U_{(R)}^2 m_i^{-1} j^{-1} = 3.25 \cdot 10^3 \text{ TeV}^{-1}$$

for 400 fb⁻¹

CMS constraints



Successful TeV-scale leptogenesis ?

$$\epsilon \gg \frac{1}{Z_f}$$

$$\epsilon = \frac{1}{H(M_1)} \frac{M_1}{m_\alpha}$$

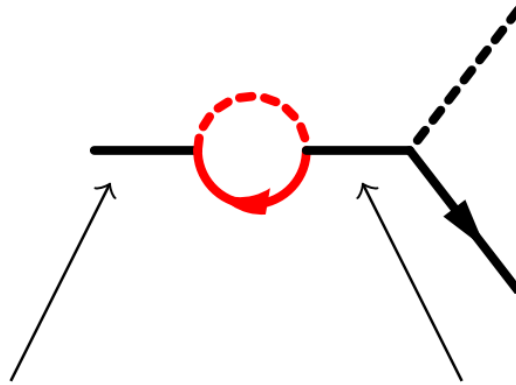
$$m_\alpha \gg 10^{13} \text{ eV}$$

$$M_1 < 5 \cdot 10^{13} \text{ eV}$$

$$\epsilon_1 \approx \frac{3}{8^{1/4}} \frac{M_1 m_3}{v^2}$$

$$M_1 \gg 10^8 \text{ GeV}$$

Resonant enhancement

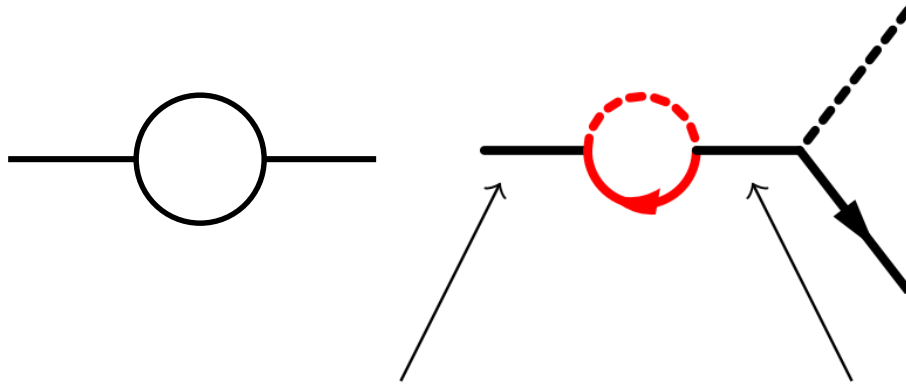


Off-shell initial N_1 : $p^2 = M_1^2 + iM_1\Gamma_1$

Internal N_2 : $\frac{i}{p^2 - M_2^2 - iM_2\Gamma_2}$

$$\epsilon_{N_i}^{wave} = \frac{\text{Im}[(h^\dagger h)_{12}^2]}{8\pi(h^\dagger h)_{ii}} \times \frac{M_1 M_2}{M_2^2 - M_1^2}$$

Schwinger-Dyson resummation

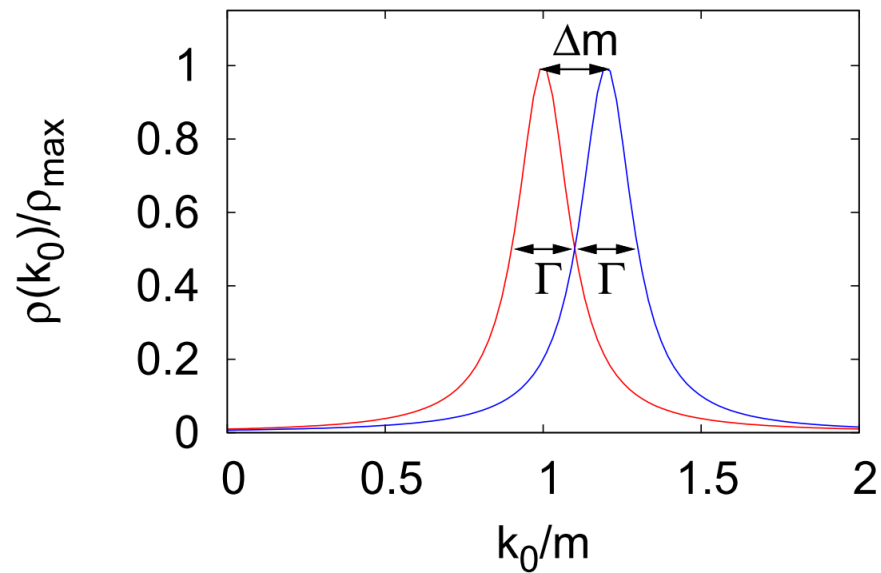
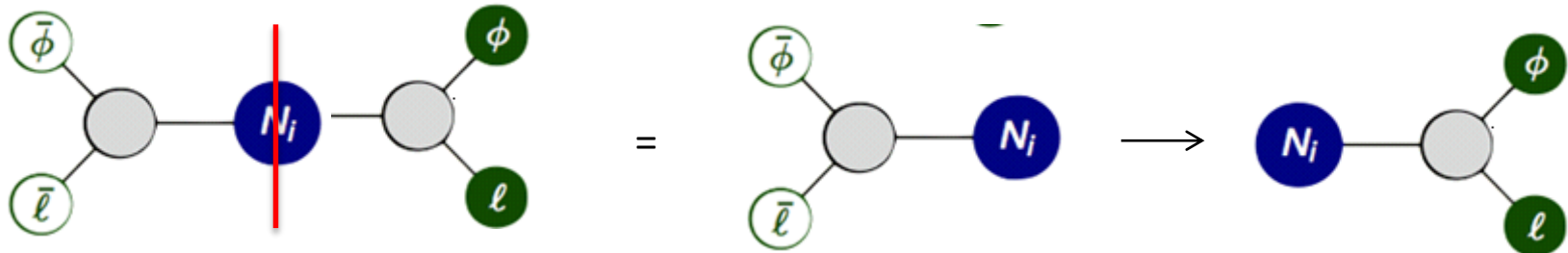


Off-shell initial N_1 : $p^2 = M_1^2 + iM_1\Gamma_1$

Internal N_2 : $\frac{i}{p^2 - M_2^2 - iM_2\Gamma_2}$

$$\epsilon_{N_i} = \frac{\text{Im}[(h^\dagger h)_{12}^2]}{8\pi(h^\dagger h)_{ii}} \frac{M_1 M_2 (M_2^2 - M_1^2)}{(M_2^2 - M_1^2 - \frac{1}{\pi} \Gamma_i M_i \ln \frac{M_2^2}{M_1^2})^2 + (M_1 \Gamma_1 - M_2 \Gamma_2)^2}$$

RIS subtraction ?



Thermal effects

$$M_i(T) = M_i(0) + \kappa_i \kappa_c (T - M_i)^2$$

$$\kappa_M(T) = \kappa_M(0) + (\kappa_2 - \kappa_1) \kappa_c (T - M_1)^2$$

$$\kappa_M \gg \kappa_1 + \kappa_2$$

$$\kappa_i / \frac{\text{Im}(H_{12}^2) (M_2^2 - M_1^2)}{(M_1^2 - M_2^2)^2 + (M_1 \kappa_1 - M_2 \kappa_2)^2}$$

Kadanoff-Baym equations

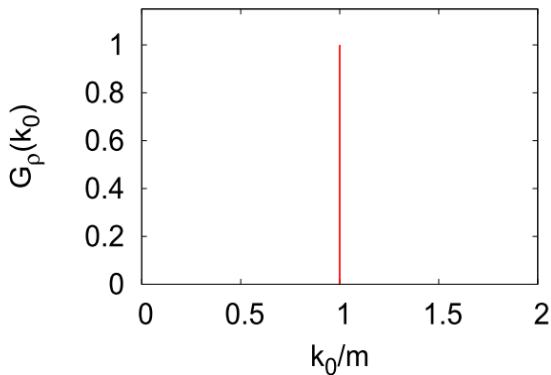
$$\begin{aligned}
 ((i\cancel{\partial}_x - M_i)\delta^{ik} - \delta\Sigma_N(x)^{ik})S_F^{kj}(x, y) &= \int_0^{x^0} dz^0 \int d^3z \Sigma_{N_\rho}^{ik}(x, z)S_F^{kj}(z, y) \\
 &\quad - \int_0^{y^0} dz^0 \int d^3z \Sigma_{N_F}^{ik}(x, z)S_\rho^{kj}(z, y) \\
 ((i\cancel{\partial}_x - M_i)\delta^{ik} - \delta\Sigma_N(x)^{ik})S_\rho^{kj}(x, y) &= \int_{y^0}^{x^0} dz^0 \int d^3z \Sigma_{N_\rho}^{ik}(x, z)S_\rho^{kj}(z, y)
 \end{aligned}$$

see hep-ph/0409233 for a comprehensive review

Quasiparticle approximation

Boltzmann limit

- ▶ on-shell quasi-stable particles

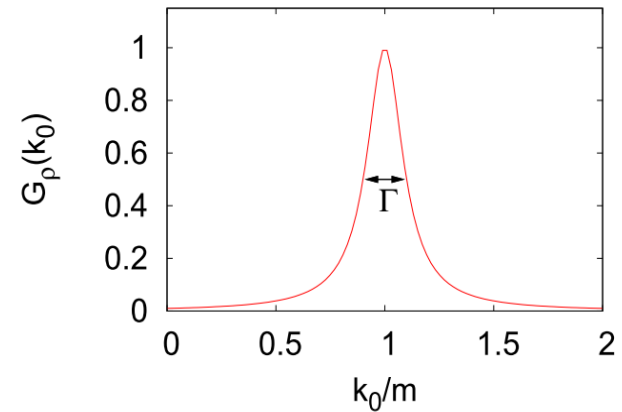


$$S_{\rho}^{ij}(k) \sim \delta^{ij} \delta(k^2 - m_i^2)$$

$$S_F^{ij}(t, k) = \left(\frac{1}{2} - f_k^i(t) \right) S_{\rho}^{ij}(k)$$

More general

- ▶ spectrum with (thermal) width



$$S_{\rho}^{ij}(t, k) \propto \frac{\delta^{ij} 2k_0 \Gamma_i(t)}{(k^2 - m_{th,i}^2(t))^2 + k_0^2 \Gamma_i(t)^2} + \dots$$

Advantages

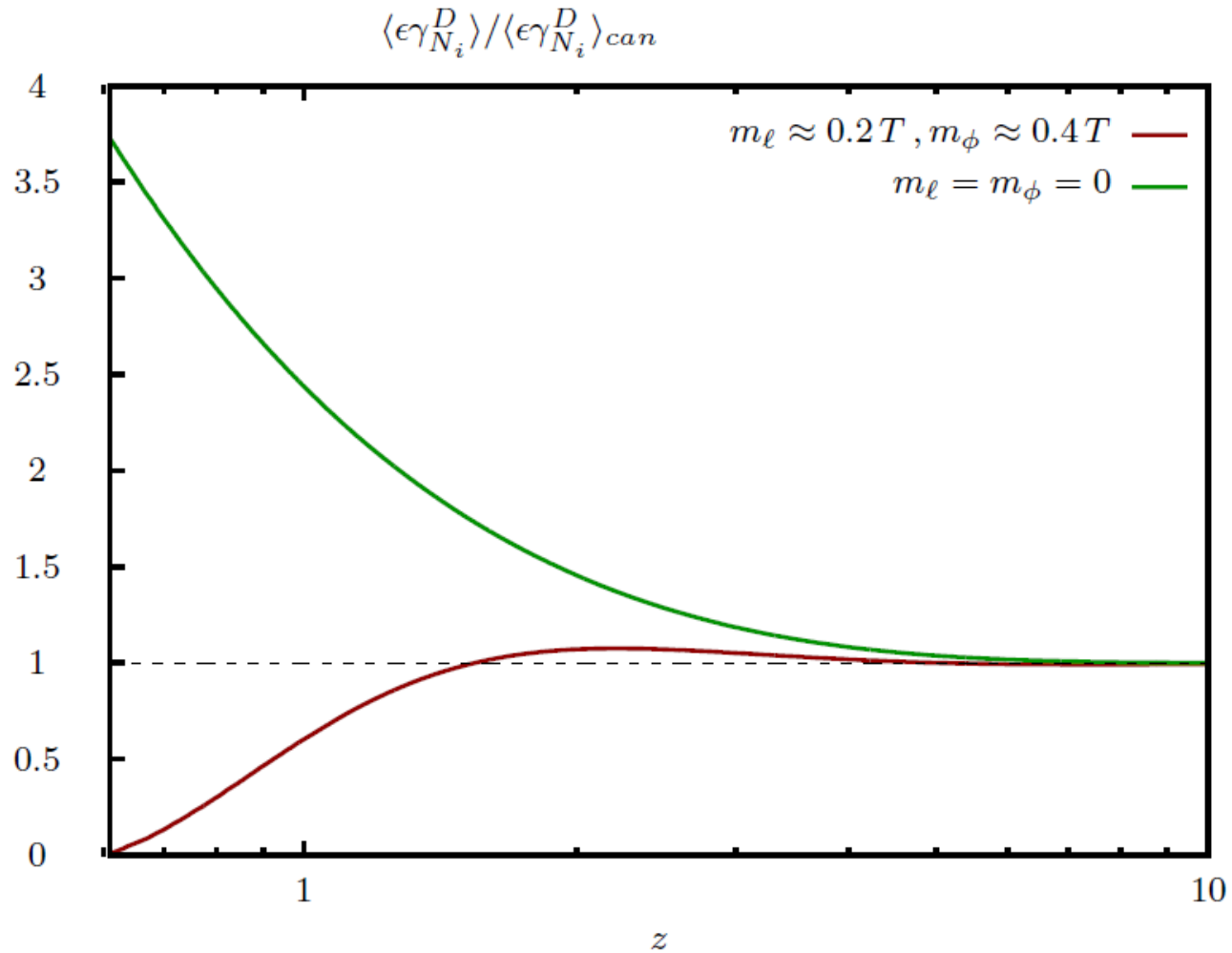
- Do not rely on quasiparticle picture
- Take medium corrections into account
- Resum series of daisy and ladder diagrams
- Take memory effects into account

Third Sakharov condition

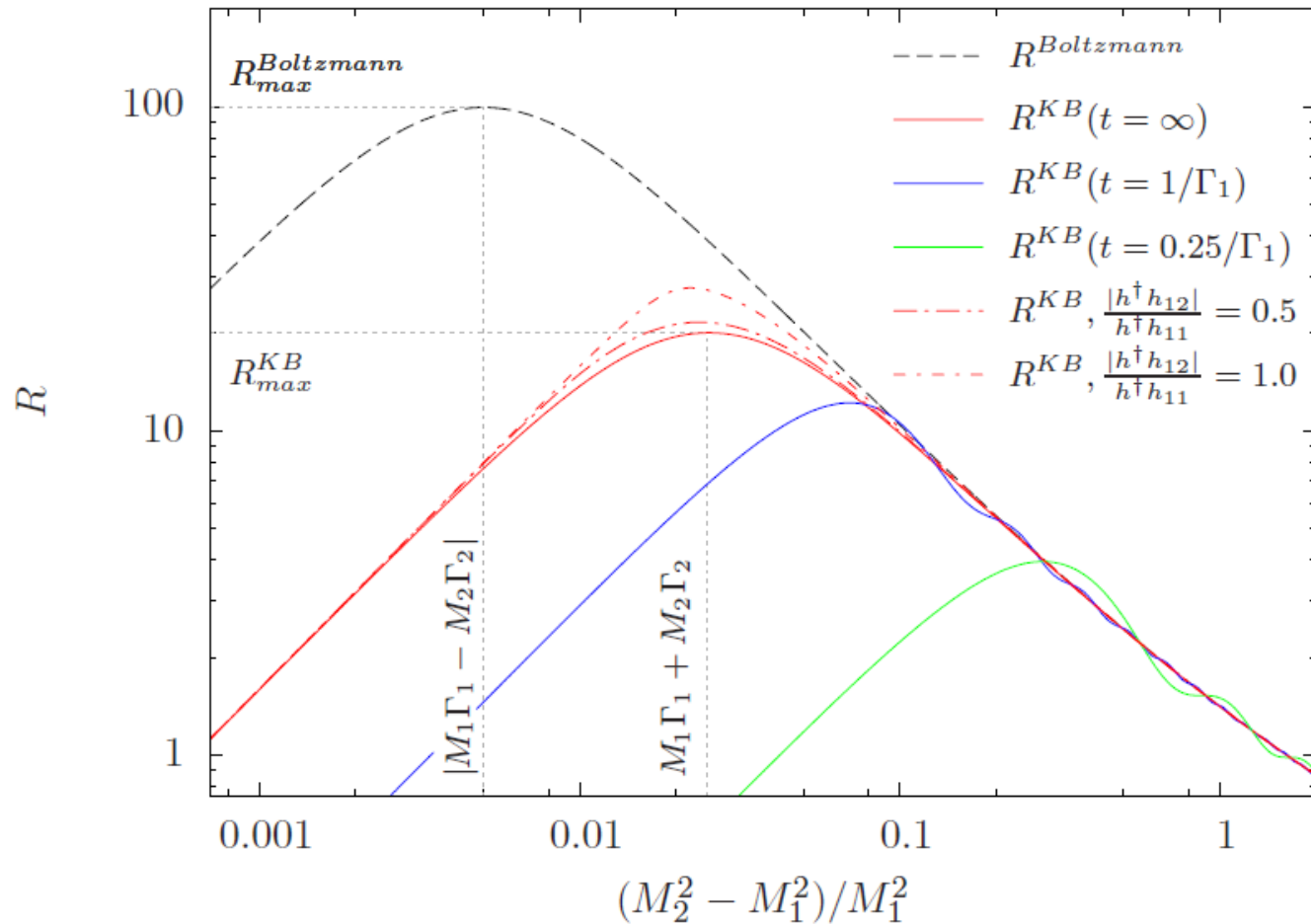
$$\begin{aligned}
 \frac{s\mathcal{H}}{z} \frac{dY_L}{dz} = & \sum_i \int d\Pi_{\ell\phi N_i}^{pkq} \mathcal{F}_{\ell\phi \leftrightarrow N_i}^{pk;q} \Xi_{\ell\phi \leftrightarrow N_i} \\
 & - \sum_i \int d\Pi_{\bar{\ell}\bar{\phi} N_i}^{pkq} \mathcal{F}_{\bar{\ell}\bar{\phi} \leftrightarrow N_i}^{pk;q} \Xi_{\bar{\ell}\bar{\phi} \leftrightarrow N_i} \\
 & - 2 \int d\Pi_{\ell\phi\bar{\ell}\bar{\phi}}^{p_1 k_1 p_2 k_2} \mathcal{F}_{\bar{\ell}\bar{\phi} \leftrightarrow \ell\phi}^{p_2 k_2; p_1 k_1} \Xi_{\bar{\ell}\bar{\phi} \leftrightarrow \ell\phi} \\
 & - \int d\Pi_{\ell\ell\bar{\phi}\bar{\phi}}^{p_1 p_2 k_1 k_2} \mathcal{F}_{\bar{\phi}\bar{\phi} \leftrightarrow \ell\ell}^{k_1 k_2; p_1 p_2} \Xi_{\bar{\phi}\bar{\phi} \leftrightarrow \ell\ell} \\
 & - \int d\Pi_{\bar{\ell}\bar{\ell}\phi\phi}^{p_1 p_2 k_1 k_2} \mathcal{F}_{\bar{\ell}\bar{\ell} \leftrightarrow \phi\phi}^{p_1 p_2; k_1 k_2} \Xi_{\bar{\ell}\bar{\ell} \leftrightarrow \phi\phi} .
 \end{aligned}$$

$$\begin{aligned}
 \mathcal{F}_{ab.. \leftrightarrow ij..}^{p_a p_b ..; p_i p_j ..} \equiv & (2\pi)^4 \delta(p_a + p_b + .. - p_i - p_j - ..) \\
 & \times [f_i f_j .. (1 - \xi^a f_a)(1 - \xi^b f_b) .. \\
 & - f_a f_b .. (1 - \xi^i f_i)(1 - \xi^j f_j) ..] ,
 \end{aligned}$$

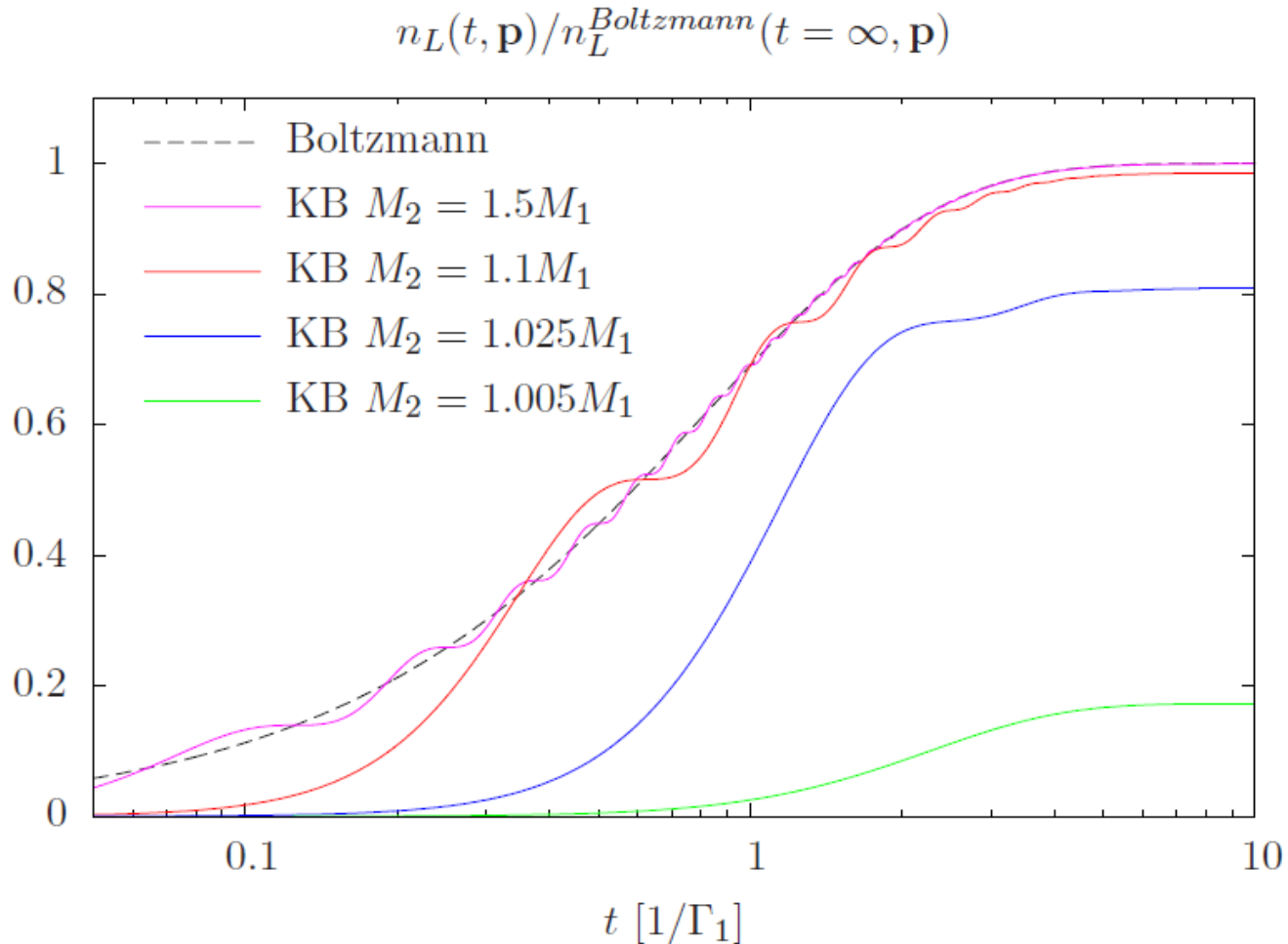
Medium enhancement of ε



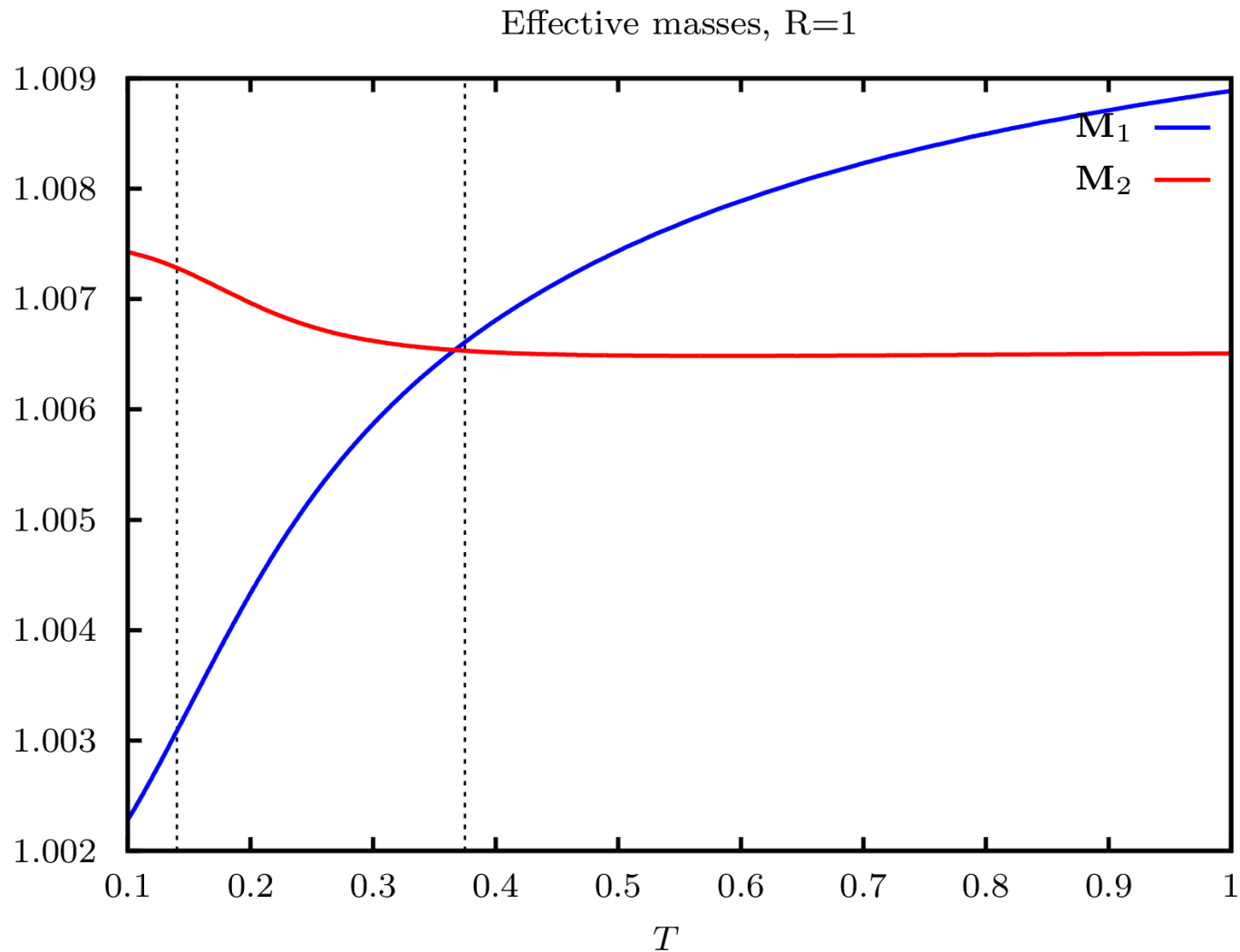
Resonant enhancement of ε



Negative interference

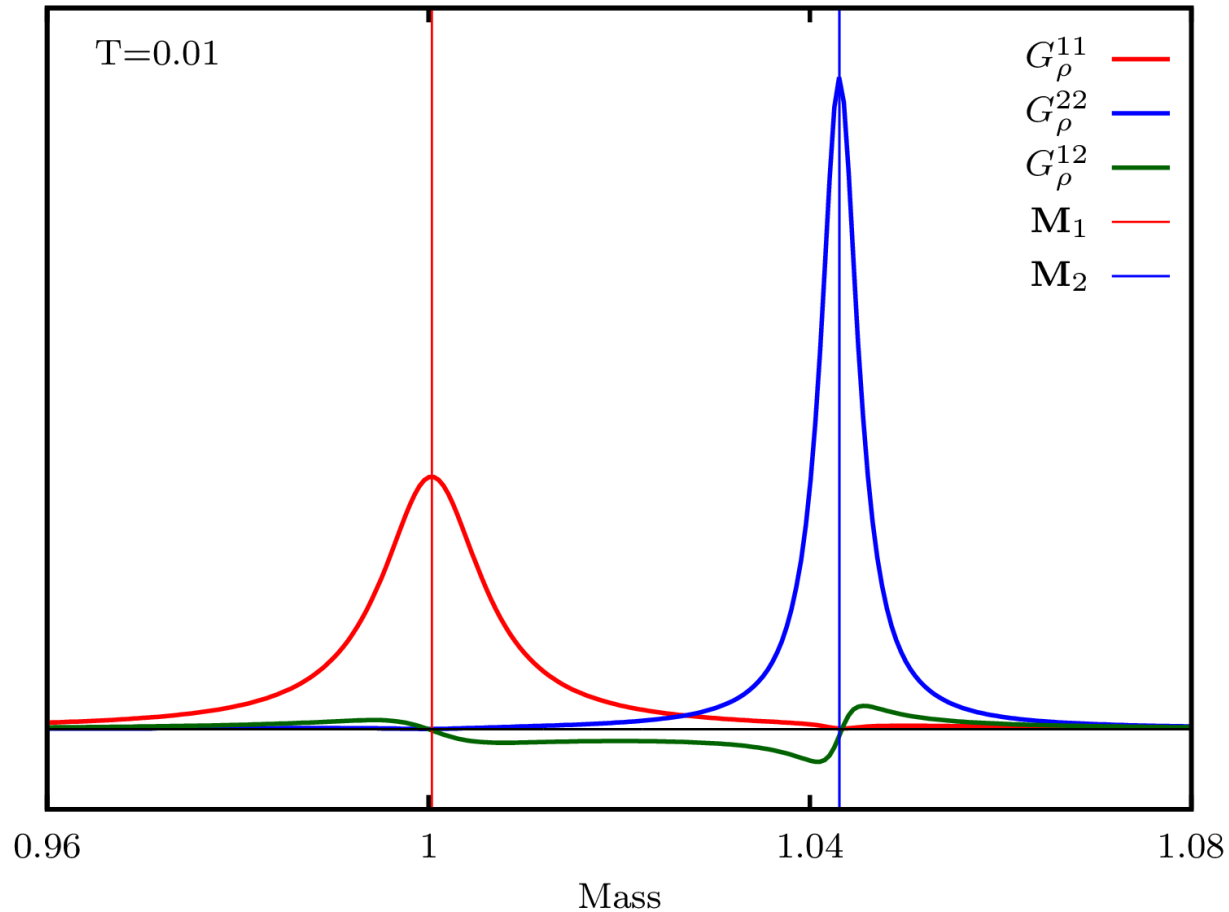


Crossing effective masses



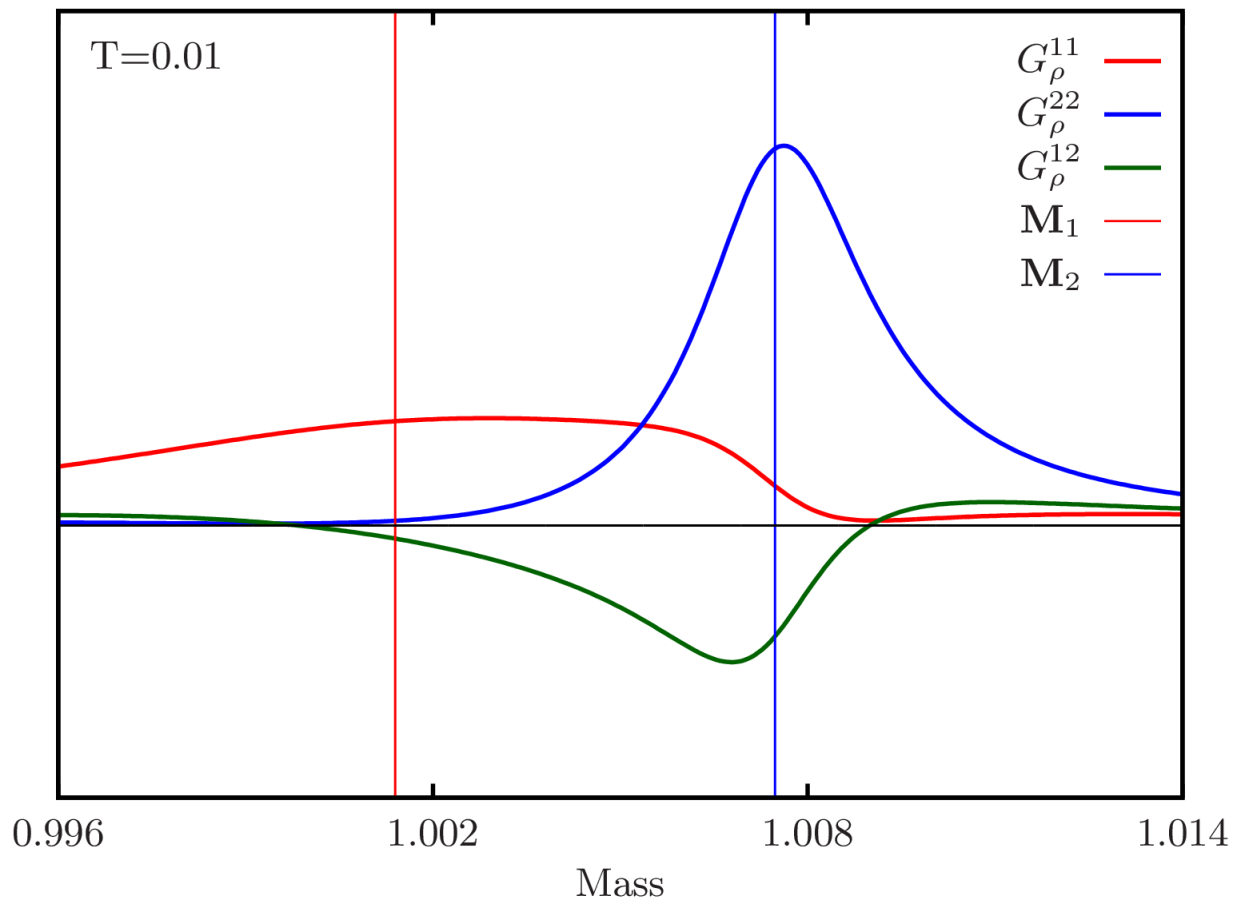
Spectral function (hierarchical)

Spectral function in crossing regime, $R=5$

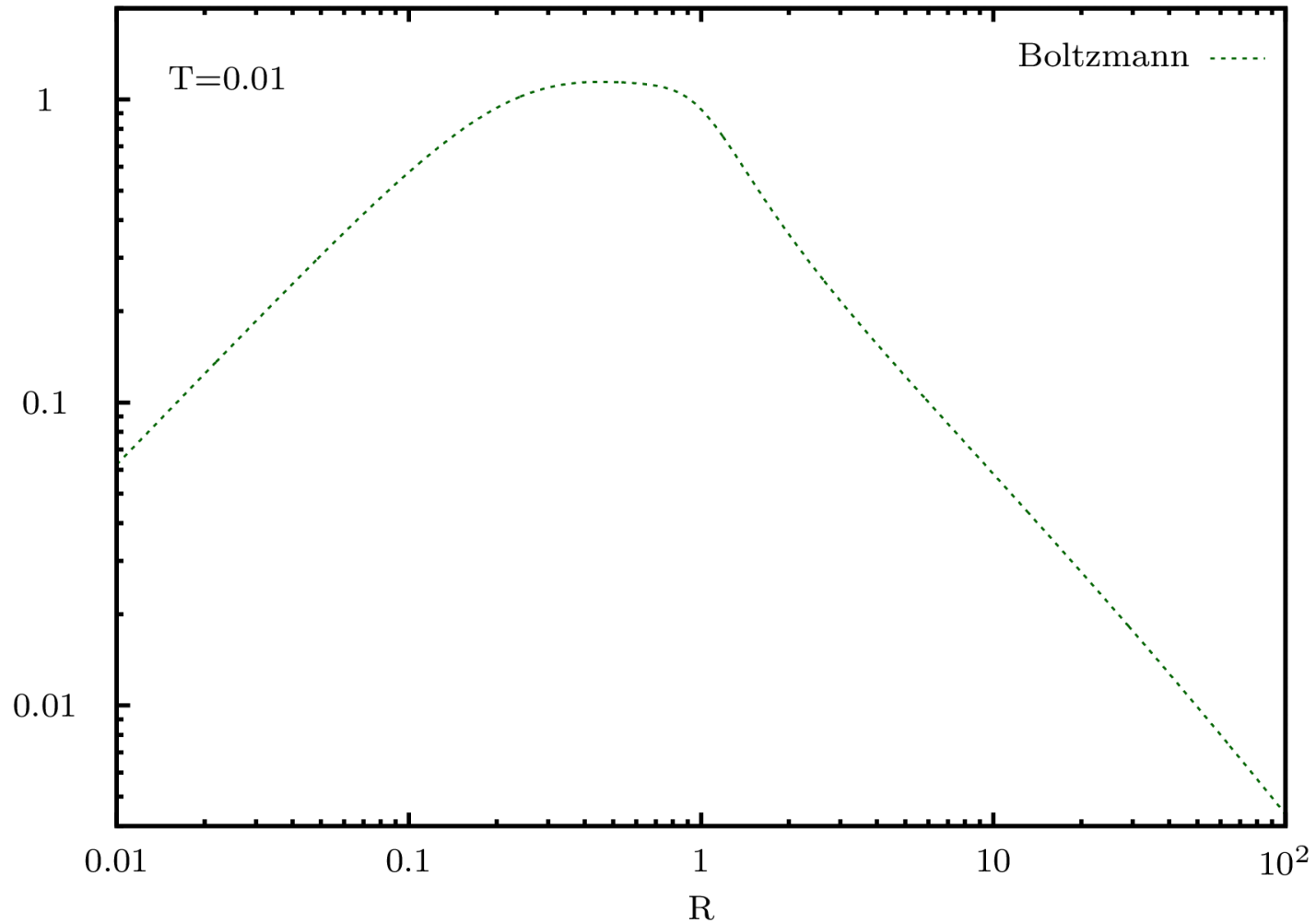


Spectral function (resonant)

Spectral function in crossing regime, $R=1$



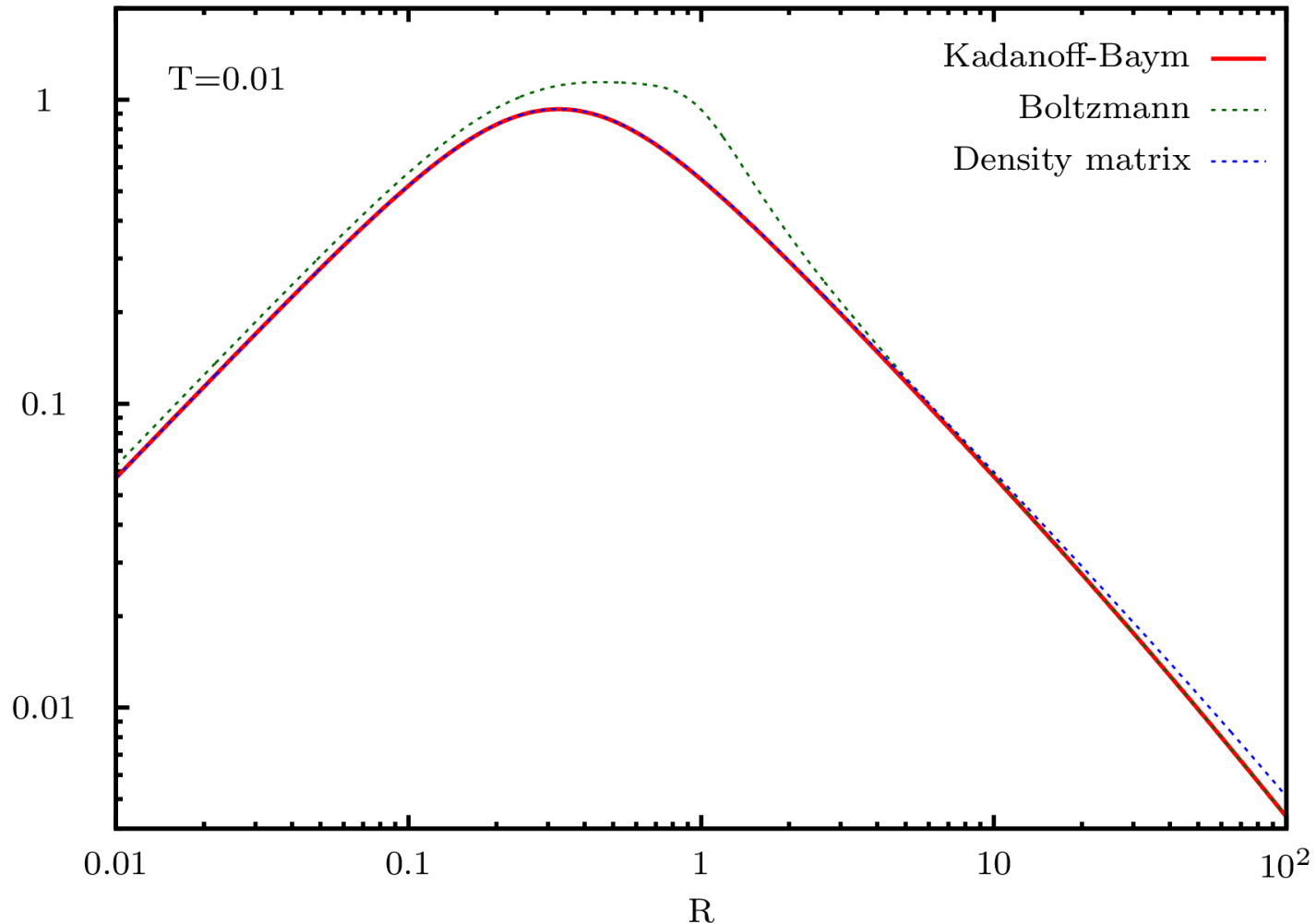
Boltzmann approximation



CP-violating parameter

$$\epsilon_2 / \frac{\text{Im}(H_{12}^2)(M_2^2 - M_1^2)}{(M_1^2 - M_2^2)^2 + (\delta_{11} - \delta_{22})^2}$$

Kadanoff-Baym result



Baryogenesis via neutrino oscillations

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(March 5, 1998)

We propose a new mechanism of leptogenesis in which the asymmetries in lepton numbers are produced through the CP-violating oscillations of “sterile” (electroweak singlet) neutrinos. The asymmetry is communicated from singlet neutrinos to ordinary leptons through their Yukawa couplings. The lepton asymmetry is then reprocessed into baryon asymmetry by electroweak sphalerons. We show that the observed value of baryon asymmetry can be generated in this way, and the masses of ordinary neutrinos induced by the seesaw mechanism are in the astrophysically and cosmologically interesting range. Except for singlet neutrinos, no physics beyond the Standard Model is required.

PACS: 98.80.Cq, 14.60.St

IC/98/22, INR-98-14T

hep-ph/9803255

1. The origin of the excess of baryons over anti-baryons in the Universe remains one of the fascinating problems of particle physics and cosmology. A number of mechanisms have been proposed to date to explain this asymmetry (for recent reviews see, *e.g.*, [1]). One of the simplest possibilities, suggested by Fukugita and Yanagida [2], is that the baryon asymmetry has originated from physics in the leptonic sector. Namely, it was assumed that at temperatures well above the electroweak scale, lepton asymmetry was produced, which was then reprocessed into the baryon asymmetry by non-perturbative electroweak effects [3] – sphalerons [4]. According to ref. [2] the lepton

neutrinos are very different from those of ref. [2].

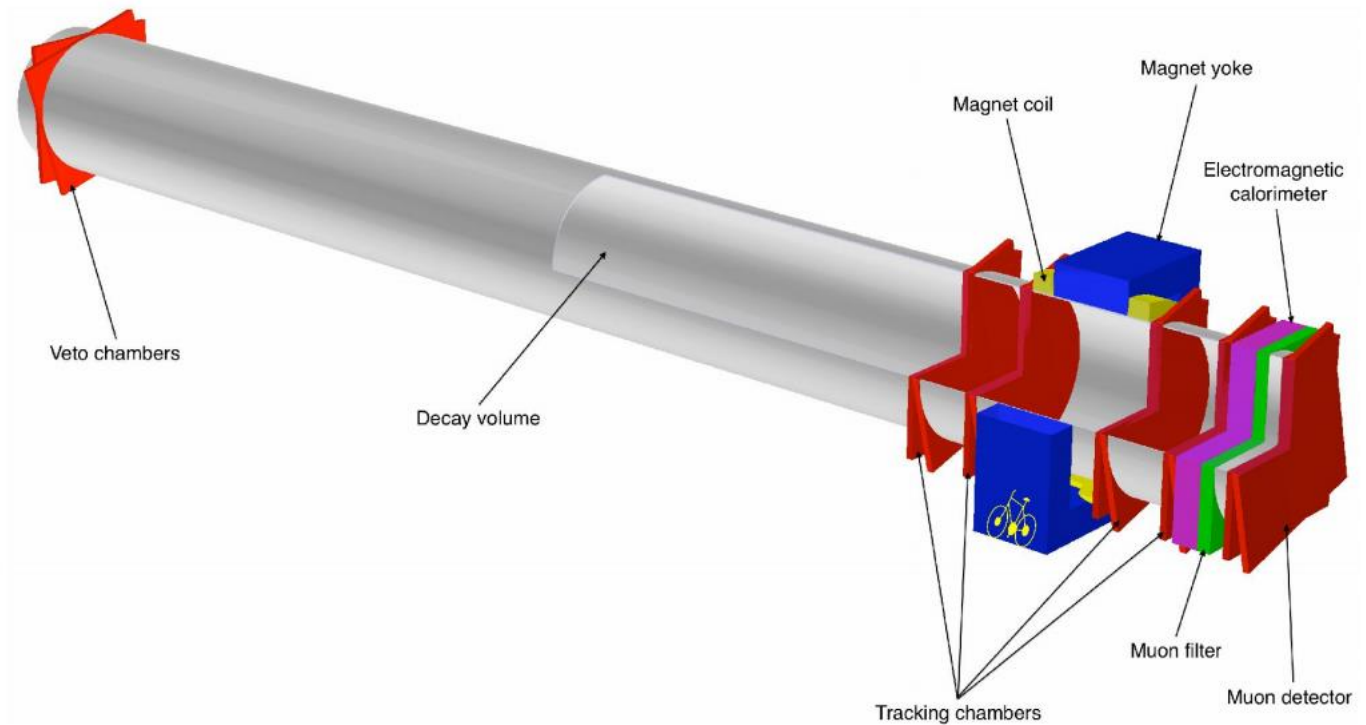
2. Let us consider the Standard Model extended by adding three types of Majorana neutrinos N_a , $a = A, B, C$ which interact with other particles only through their Yukawa couplings [6]. The corresponding Lagrangian can be written in the “Yukawa basis” (where the matrix of Yukawa coupling constants has been diagonalized) as follows,

$$\mathcal{L} = \bar{N}_{Ra} i \not{\partial} N_{Ra} + h_a \bar{l}_a N_{Ra} \Phi + \frac{M_{ab}}{2} N_{Ra}^T C N_{Rb} + h.c. .$$

ν MSM

Quarks	2.4 MeV $\frac{2}{3}$ Left u up Right	1.27 GeV $\frac{2}{3}$ Left c charm Right	171.2 GeV $\frac{2}{3}$ Left t top Right		
	4.8 MeV $-\frac{1}{3}$ Left d down Right	104 MeV $-\frac{1}{3}$ Left s strange Right	4.2 GeV $-\frac{1}{3}$ Left b bottom Right		
	<0.0001 eV 0 Left ν_e electron neutrino Right	$\sim \text{keV}$ N_1 sterile neutrino	$\sim 0.01 \text{ eV}$ ν_μ muon neutrino	$\sim \text{GeV}$ N_2 sterile neutrino	$\sim 0.04 \text{ eV}$ ν_τ tau neutrino
Leptons	0.511 MeV -1 Left e electron Right	105.7 MeV -1 Left μ muon Right	1.777 GeV -1 Left τ tau Right		

SHIP



Summary

- Majorana neutrinos solve two problems
- TeV-scale Majoranas are searched for at LHC
- GeV-scale Majoranas are searched for at SHIP
- In both cases need resonant leptogenesis

Idea of SHIP

