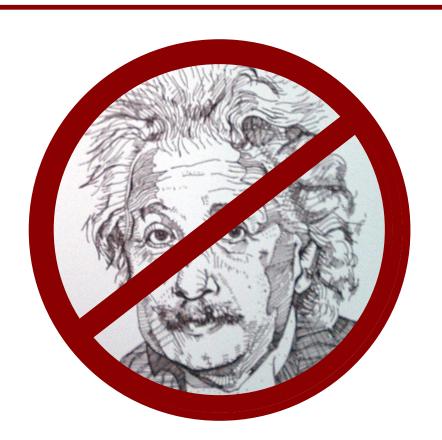
# Nonrelativistic, non-Newtonian gravity





Dieter Van den Bleeken Boğaziçi University

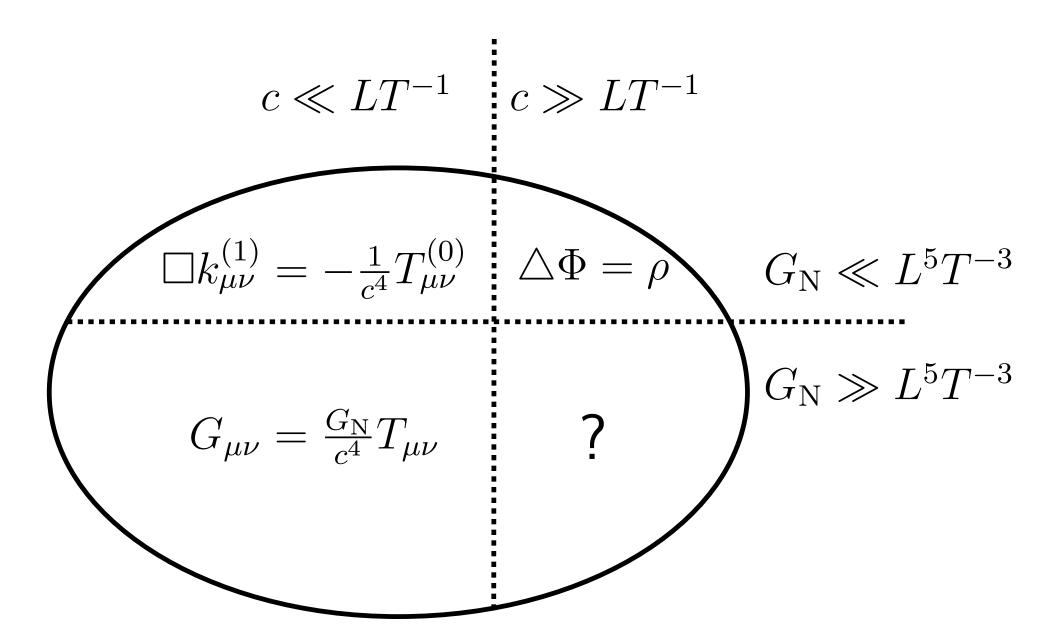
based on arXiv:1512.03799 and work in progress with Çağın Yunus

IPM Tehran

27<sup>th</sup> May 2016

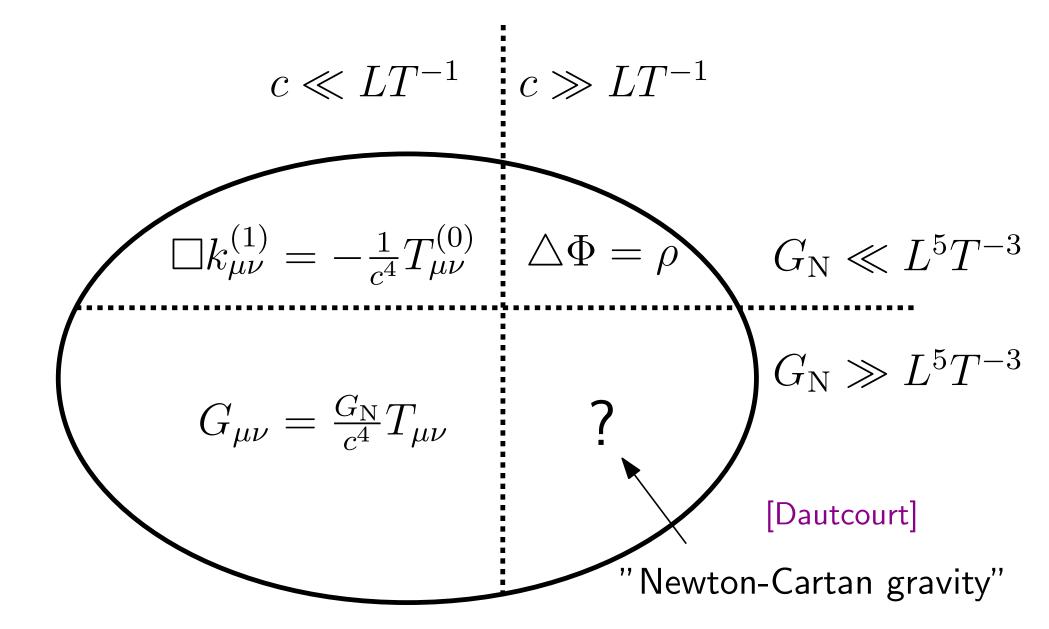
# Nonrelativistic, non-Newtonian gravity

## Gravity in various regimes



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# Gravity in various regimes



#### Outline

- Motivation
- From NG towards GR
- NC gravity/geometry
- From GR to NC
- $NC \supseteq NG$ 
  - → dimensional reduction in NC
- From GR to TTNC
- Example

## Legend

NG = Newtonian Gravity GR = General Relativity

NC = Newton-Cartan TT = Twistless Torsional

#### Outline

- Motivation
- From NG towards GR
- NC gravity/geometry
- From GR to NC
- NC ⊋ NG
   → dimensional reduction in NC
- From GR to TTNC
- Example

[Cartan; Trautman; Kunzl]
Review
[Dautcourt; Tichy, Flanagan]

Related work
[Afshar, Bergshoeff,
Hartong et al.]

New

Legend

 $NG = Newtonian Gravity \qquad GR = General Relativity$  $<math>NC = Newton-Cartan \qquad TT = Twistless Torsional$ 

### Motivation

- Isn't GR our best theory of gravity?
  - → our most complicated theory of gravity all aproximation schemes are welcome

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  - \* condensed matter on non-flat backgrounds
  - \* non-relativistic holography

#### Motivation

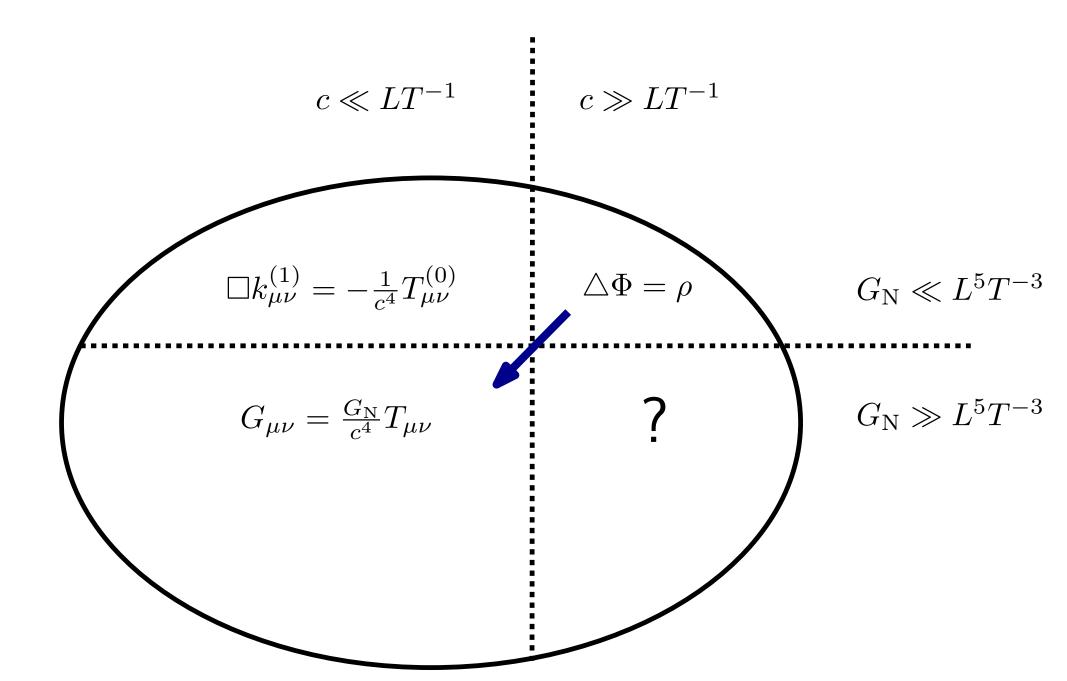
- Isn't GR our best theory of gravity?
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- Why nonrelativistic?
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  - \* condensed matter on non-flat backgrounds
  - \* non-relativistic holography
- Why non-Newtonian?
  - \* analytic description of strong gravity!

isn't this an empty set? actually, no...

is this a physically relevant regime?

wasn't this done 50 years ago? please let me know!

## From NG towards GR



### From NG towards GR

Geometrize Newtonian gravity

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Geometrize Newtonian gravity

$$ec{a}=-ec{
abla}\Phi$$
  $\Leftrightarrow$  geodesic equation 
$$\begin{picture}(1,0) \put(0,0){\line(0,0){100}} \put(0$$

⇒ covariantize this: Newton-Cartan geometry/gravity

## Newton-Cartan geometry/gravity

Geometry

fields: 
$$au_{\mu}\,, \quad h^{\mu 
u}\,, \quad \Gamma^{\lambda}_{\mu 
u}$$

constraints:

$$\tau_{\mu}h^{\mu\nu} = 0, \quad h^{[\mu\nu]} = 0, \quad \Gamma^{\lambda}_{[\mu\nu]} = 0$$

$$\nabla_{\mu}\tau_{\nu} = 0 \qquad \nabla_{\mu}h^{\nu\lambda} = 0$$

$$h^{\lambda(\mu}R^{\nu)}{}_{\lambda\rho\sigma}(\Gamma) = 0$$

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Gravity

EOM:  $R_{\mu\nu}(\Gamma) = \tau_{\mu}\tau_{\nu}\,\rho$ 

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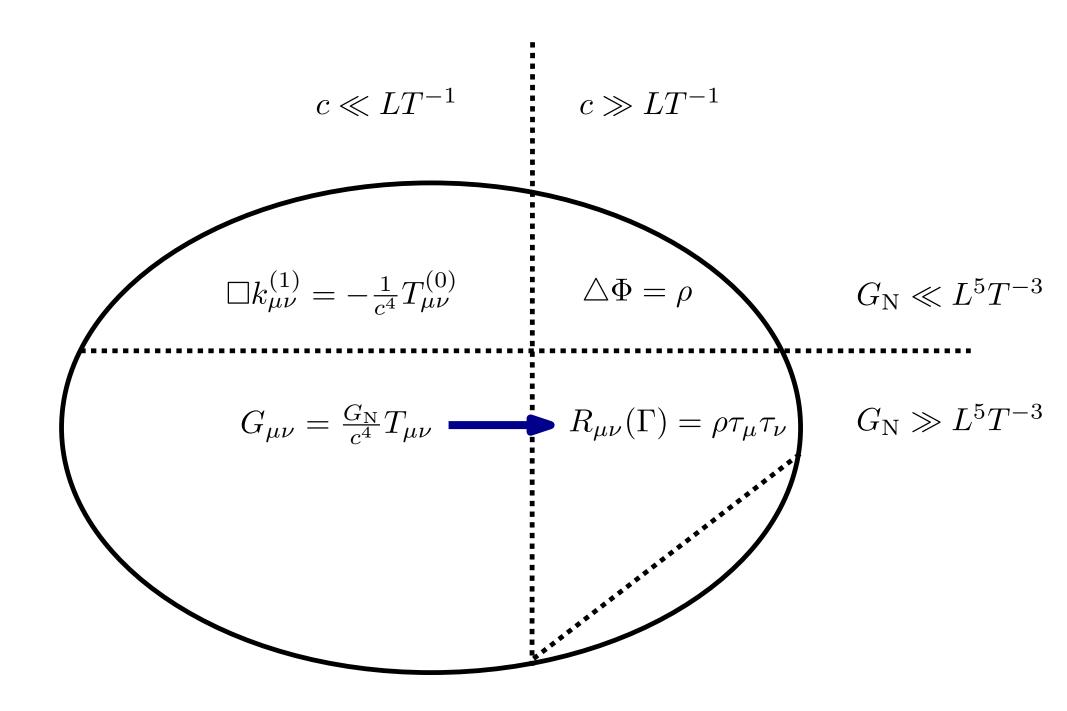
Gravity

EOM:  $R_{\mu\nu}(\Gamma) = \tau_{\mu}\tau_{\nu} \rho$ 

Manifest 1+3 dim coordinate invariance

Equal to Newtonian gravity when

$$\tau_{\mu} = \delta_{\mu}^{0}, \ h^{\mu 0} = 0, \ h^{ij} = \delta^{ij}, \ \Gamma_{0i}^{j} = 0$$



ullet Large c expansion

$$g_{\mu\nu}(c) = \sum_{i=-1}^{\infty} \overset{(2i)}{g}_{\mu\nu} c^{-2i} \qquad g^{\mu\nu}(c) = \sum_{i=0}^{\infty} \overset{(2i)}{g}^{\mu\nu} c^{-2i}$$
 Ansatz:  $\overset{(-2)}{g}_{\mu\nu} = -\tau_{\mu}\tau_{\nu}$  (natural via  $x^0 = ct$ )

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Solving invertibility leads to

$$\begin{array}{l} \overset{(0)}{g}{}^{\mu\nu} = h^{\mu\nu} \\ \overset{(0)}{g}{}_{\mu\nu} = 2\tau_{(\mu}C_{\nu)} + h_{\mu\nu} \qquad \overset{(2)}{g}{}^{\mu\nu} = -\tau^{\mu}\tau^{\nu} + 2\tau^{(\mu}h^{\nu)\lambda}C_{\lambda} \\ \text{where} \qquad \tau^{\mu}\tau_{\nu} + h^{\mu\lambda}h_{\lambda\nu} = \delta^{\mu}_{\nu} \end{array}$$

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Metric compatibility & finite Levi-Civita implies that

$$\nabla_{\mu}\tau_{\nu} = 0 \qquad \nabla_{\mu}h^{\nu\lambda} = 0 \qquad K_{\mu\nu} = 2\partial_{[\mu}C_{\nu]}$$
$$\Gamma^{\lambda}_{\mu\nu} = \frac{1}{2}h^{\lambda\rho}\left(\partial_{\mu}h_{\rho\nu} + \partial_{\nu}h_{\mu\rho} - \partial_{\rho}h_{\mu\nu}\right) + \tau^{\lambda}\partial_{(\mu}\tau_{\nu)} + h^{\lambda\rho}\tau_{(\mu}K_{\nu)\rho}$$

Expand Einstein equations

(for perfect fluid)

$$\overset{\mbox{\tiny (-4)}}{R}_{\mu\nu} = \overset{\mbox{\tiny (-2)}}{R}_{\mu\nu} = 0$$
 — automatic

$$R_{\mu\nu}^{(0)} = \mathcal{T}_{\mu\nu} \quad \Leftrightarrow \quad [R_{\mu\nu}(\Gamma) = \rho \, \tau_{\mu} \tau_{\nu}]$$

$$c \ll LT^{-1}$$

$$c \gg LT^{-1}$$

$$\Box k_{\mu\nu}^{(1)} = -\frac{1}{c^4} T_{\mu\nu}^{(0)}$$

$$G_{\mu\nu} = \frac{G_{\rm N}}{c^4} T_{\mu\nu}$$

$$\triangle \Phi = \rho$$

$$R_{\mu\nu}(\Gamma) = \rho \tau_{\mu} \tau_{\nu}$$

$$G_{\rm N} \ll L^5 T^{-3}$$

$$G_{\rm N}\gg L^5T^{-3}$$

The dof's of NC manifest themselves after partial gauge-fixing  ${\sf Diff}_4 \to {\sf Diff}_3(t)$ 

$$R_{ij} = 0$$

$$-\nabla^j K_{ji} = 2h^{jk} \nabla_{[i} \dot{h}_{j]k}$$

$$-\nabla^i G_i = \frac{1}{2} h^{ij} \ddot{h}_{ij} + \frac{1}{4} \dot{h}_{ij} \dot{h}^{ij} - \frac{1}{4} K_{ij} K^{ij} + 4\pi \rho$$
where  $G_i = -\partial_i \Phi - \dot{C}_i$   $\Phi = -C_0$ 

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## Standard argument:

- 1) In 3d Ricci flat = flat
- 2)  $C_i$  is harmonic, hence constant
- 3) NC=NG

Conclusion: nonrelativistic non-Newtonian gravity doesn't exist

The dof's of NC manifest themselves after partial gauge-fixing  ${\sf Diff}_4 o {\sf Diff}_3(t)$ 

$$\begin{array}{rcl} R_{ij}&=&0\\ -\nabla^j K_{ji}&=&2h^{jk}\nabla_{[i}\dot{h}_{j]k}\\ -\nabla^i G_i&=&\frac{1}{2}h^{ij}\ddot{h}_{ij}+\frac{1}{4}\dot{h}_{ij}\dot{h}^{ij}-\frac{1}{4}K_{ij}K^{ij}+4\pi\rho\\ \end{array}$$
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Standard argument:

What about other energy-momentum?

- 1) In 3d Ricci flat  $\longrightarrow$  What about higher dimensions?
- 2)  $C_i$  is harmonic, hence constant
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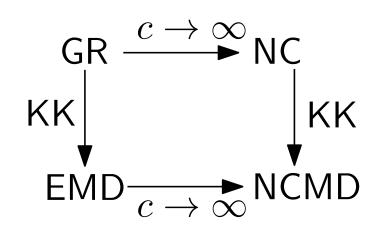
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dimensional reduction

# Newton - Cartan $\supseteq$ Newtonian Gravity

#### dimensional reduction

We showed



[arXiv: 1512.03799]

Here is NCMD

$$\nabla_{i}\partial^{i}\Omega = \frac{1}{4}\Omega^{3}F_{ij}F^{ij} \qquad \nabla_{i}\left(\Omega^{3}E^{i}\right) = \frac{1}{2}\Omega^{3}K_{ij}F^{ij} \qquad \nabla_{i}\left(\Omega^{3}F^{ij}\right) = 0$$

$$\Omega R_{ij} = \nabla_{i}\partial_{j}\Omega + \frac{1}{2}\Omega^{3}F_{i}{}^{k}F_{jk}$$

$$\nabla_{j}\left(\Omega K^{j}{}_{i}\right) = -\Omega\partial_{i}(h^{jk}\dot{h}_{jk}) + \nabla^{j}(\Omega\dot{h}_{ij}) - 2\partial_{i}\dot{\Omega} + \Omega^{3}F_{ij}E^{j}$$

$$-\nabla_{i}\left(\Omega G^{i}\right) = \frac{1}{4}\Omega\left(\dot{h}^{ij}\dot{h}_{ij} - K_{ij}K^{ij} + 2h^{ij}\ddot{h}_{ij}\right) + \frac{1}{2}\Omega^{3}E_{i}E^{i} + \ddot{\Omega}$$



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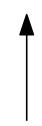
- Large c expansion  $\checkmark$
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Torsional Newton - Cartan geometry

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Expanding Einstein equations

$$R_{\mu\nu}^{\text{\tiny (-4)}} = 0 \quad \Leftrightarrow \quad \tau_{[\mu}\partial_{\nu}\tau_{\lambda]} = 0 \quad \Leftrightarrow \quad \partial_{[\mu}\tau_{\nu]} = \tau_{[\mu}a_{\nu]}$$

Twistless Torsional NC geometry

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$$R_{\mu\nu}^{\text{\tiny (-2)}} = \tau_{\mu}\tau_{\nu}D_{\rho}a^{\rho} \qquad D_{\mu} = \nabla_{\mu} - a_{\mu}$$

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$$+\tau^{\lambda}\partial_{[\mu}\tau_{\nu]} + h^{\lambda\rho}\left(C_{\mu}\partial_{[\nu}\tau_{\rho]} + C_{\nu}\partial_{[\mu}\tau_{\rho]} - C_{\rho}\partial_{[\mu}\tau_{\nu]}\right)$$

Expanding Einstein equations

$$\begin{split} R_{\mu\nu}^{^{(-4)}} &= 0 \quad \Leftrightarrow \quad \tau_{[\mu} \partial_{\nu} \tau_{\lambda]} = 0 \quad \Leftrightarrow \quad \partial_{[\mu} \tau_{\nu]} = \tau_{[\mu} a_{\nu]} \\ R_{\mu\nu}^{^{(-2)}} &= \tau_{\mu} \tau_{\nu} \, D_{\rho} a^{\rho} \qquad \qquad D_{\mu} = \overset{^{(\text{nc})}}{\nabla}_{\mu} - a_{\mu} \\ R_{\mu\nu}^{^{(0)}} &= \overset{^{(\text{nc})}}{R}_{\mu\nu} + D^{\rho} \Big( \hat{h}_{\rho\nu} \hat{a}_{\mu} \Big) - \frac{1}{2} a^{\rho} \overset{^{(\text{nc})}}{\nabla}_{\rho} \hat{h}_{\mu\nu} - a_{\rho} \tau_{(\mu} \overset{^{(\text{nc})}}{\nabla}_{\nu)} \hat{\tau}^{\rho} + \frac{1}{2} \tau_{\mu} \tau_{\nu} a^{2} \hat{\tau}^{2} \end{split}$$

Schwarzschild

$$ds^{2} = -c^{2} \left( 1 - \frac{2m}{c^{2} r} \right) dt^{2} + \left( 1 - \frac{2m}{c^{2} r} \right)^{-1} dr^{2} + r^{2} d\Omega^{2}$$
$$= -c^{2} dt^{2} + \frac{2m}{r} dt^{2} + dr^{2} + r^{2} d\Omega^{2} + \mathcal{O}(c^{-2})$$

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$$\left(\tau_0 = 1, \ \tau_i = 0 \qquad h_{ij} = \delta_{ij}, \ h_{\mu 0} = 0 \qquad C_0 = \frac{m}{r}, \ C_i = 0\right)$$

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$$\left[\tau_0 = 1, \ \tau_i = 0 \qquad h_{ij} = \delta_{ij}, \ h_{\mu 0} = 0 \qquad C_0 = \frac{m}{r}, \ C_i = 0\right]$$

- This solves NC eom
- $\Phi = C_0 = \frac{m}{r}$  Newtonian gravity of point mass
- $d\tau = 0 \Rightarrow a_{\mu} = 0$  no torsion

Schwarzschild (extremely massive)

$$ds^{2} = -c^{2} \left( 1 - \frac{2M}{r} \right) dt^{2} + \left( 1 - \frac{2M}{r} \right)^{-1} dr^{2} + r^{2} d\Omega^{2}$$

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$$\tau_{\mu} = \left(1 - \frac{2M}{r}\right)^{1/2} \delta_{\mu}^{t}, \quad h_{\mu\nu} = \begin{pmatrix} 0 & 0 & 0 & 0 & 0 \\ 0 & \left(1 - \frac{2M}{r}\right)^{-1} & 0 & 0 \\ 0 & 0 & r^{2} & 0 \\ 0 & 0 & 0 & r^{2} \sin^{2} \theta \end{pmatrix} \quad C_{\mu} = 0$$

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- This solves TTNC eom
- $\Phi = C_0 = 0$  vanishing Newtonian potential!
- $d\tau \neq 0 \Rightarrow a_{\mu} = -\frac{M}{r^2} \left(1 \frac{2M}{r}\right)^{-1} \delta_{\mu}^{r}$  non-vanishing torsion!
- curved spatial geometry!

Nonrelativistic non-Newtonian gravity

# Summary & outlook

# Two punch lines:

- nonrelativistic gravity is more than a Newtonian potential
- Appearance of TTNC out of GR

# In progress/to do:

- precise relation with bottom up TTNC
- add energy momentum
- gauge fixed version
- precise physical interpretation
- real world applications?

#### Gaussian normal coordinates

There always exist coordinates such that (on a patch)

$$ds^2 = -c^2 d\sigma^2 + g_{ij} dx^i dx^j$$

⇒ doesn't that imply we can always remove the torsion?

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⇒ doesn't that imply we can always remove the torsion?

The transformation to GN coordinates is not compatible with (standard) large c expansion

Schwarzschild (extremely massive)

$$ds^{2} = -c^{2}d\sigma^{2} + \frac{2M}{r(\sigma,\rho)}d\rho^{2} + r(\sigma,\rho)^{2}d\Omega^{2}$$
$$= -c^{2}d\sigma^{2} + c^{4/3}\left(\frac{9}{2}M\right)^{2/3}\sigma^{4/3}d\Omega^{2} + \mathcal{O}(c^{1/3})$$

$$r(\sigma, \rho) = \left(-\frac{3}{2}\sqrt{2M}(c\sigma + \rho)\right)^{2/3}$$