

Real time dynamics and phase separation in a holographic 1st order phase transition

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Physics, IPM

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Janik, Jankowski, HS, 1704.05387
Bajnok, Janik, Jankowski, HS (in progress)

Simple exercise (homework):

$$rS((\partial_r G)^2 + G^2(\partial_r \Phi)^2) + 4G^2(2\partial_r S + r\partial_r^2 S) = 0,$$

$$G^2 S(2r^4 \partial_r A \partial_r S + S(r^4 \partial_r^2 A + 2r^3 \partial_r A - r^2 \partial_t \Phi \partial_r \Phi + V(\Phi)) - 4r^2 \partial_r \partial_t S) \\ - r^2 (\partial_r G (r^2 B \partial_r B + \partial_t G S^2) - \partial_r B \partial_x G) + r^2 (r^2 \partial_r B^2 - \partial_r \partial_x B + rB(2\partial_r B + r\partial_r^2 B))G = 0,$$

$$-GS(S(r^2 \partial_r B \partial_r G - \partial_r \partial_x G) + rB(2r\partial_r G \partial_r S + (2\partial_r G + r\partial_r^2 G)S) - 2\partial_r G \partial_x S) \\ + G^2(S^2(r^2 \partial_r^2 B + 2r\partial_r B - \partial_x \Phi \partial_r \Phi) + 2\partial_r S(\partial_x S - r^2 B \partial_r S) - 2S(rB(2\partial_r S + r\partial_r^2 S) + \partial_r \partial_x S)) \\ + \partial_r G S^2(r^2 B \partial_r G - 2\partial_x G) = 0,$$

$$\partial_x A(r^2(\partial_r B G + B \partial_r G) - \partial_x G) - r^2 \partial_r A \partial_x B G + AG(GS(2r^4 \partial_r A \partial_r S + S(r^4 \partial_r^2 A + 2r^3 \partial_r A + V(\Phi)) \\ + r^4 \partial_r B^2) - r^4 \partial_r A B^2 \partial_r G + r^2 \partial_r A B \partial_x G + r^4 \partial_r^2 A B^2 G + 2r^3 \partial_r A B^2 G - 2r^2 \partial_r \partial_x A B G \\ - 2r^2 \partial_r A G^2 S \partial_t S + 2r^2 \partial_t A G^2 S \partial_r S + \partial_t^2 A G - 2r^2 B \partial_t B \partial_r G + 2\partial_t B \partial_x G + 2r^2 B \partial_r \partial_t B G - 2\partial_t \partial_x B G \\ - \partial_t G^2 S^2 - 4G^2 S S^{(0,2,0)}(r, t, x) - G^2 S^2 \partial_t \Phi^2 = 0,$$

$$S^2(G(r^2 \partial_r G(-\partial_x A + r^2 A \partial_r B - \partial_t B)) + G(r^2(\partial_r \partial_x A + rA(2\partial_r B + r\partial_r^2 B) - \partial_r \partial_t B) - \partial_x \Phi \partial_t \Phi) \\ + \partial_t \partial_x G - 2\partial_x G \partial_t G) + B(r^2 G(S^2(r^2 \partial_r A \partial_r G - \partial_r \partial_t G) + r^2 \partial_r B^2 - \partial_r \partial_x B - 2\partial_t G \partial_r S S) \\ + G^2(2S(r^4 \partial_r A \partial_r S - r^2 \partial_r \partial_t S) - 2r^2 \partial_t S \partial_r S + S^2 V(\Phi)) + r^2(\partial_r B \partial_x G + \partial_t G \partial_r G S^2)) + 2G^2 \partial_x S \partial_t S \\ + r^3 B^2((2\partial_r B + r\partial_r^2 B)G - r\partial_r B \partial_r G) - 2GS(G(r^2(\partial_r B \partial_t S - \partial_t B \partial_r S) + \partial_t \partial_x S) - \partial_t G \partial_x S) = 0,$$

$$G^2(-r^2 S^4(r^2 \partial_r A \partial_r G + rA(2\partial_r G + r\partial_r^2 G) - 2\partial_r \partial_t G) + 2r^2 S^3(\partial_r S(\partial_t G - r^2 A \partial_r G) + \partial_r G \partial_t S) \\ - 2r^2 \partial_r B S \partial_x S - 2S(r^2(rB^2(2\partial_r S + r\partial_r^2 S) - \partial_x B \partial_r S) + \partial_x^2 S) - S^2(r^4 \partial_r B^2 - 2r^2 \partial_r \partial_x B \\ + \partial_x \Phi^2) - 2r^4 B^2 \partial_r S^2 + 2\partial_x S^2) + G^3(-S^2)(2r^2 S(r^2 \partial_r A \partial_r S + rA(2\partial_r S + r\partial_r^2 S) - 2\partial_r \partial_t S) \\ + 2r^2 \partial_r S(r^2 A \partial_r S - 2\partial_t S) + S^2 V(\Phi)) + GS(r^2 \partial_r G S^3(r^2 A \partial_r G - 2\partial_t G) + \partial_x G(2\partial_x S - r^2(\partial_r B S \\ + 2B \partial_r S)) + S(\partial_x^2 G - r^2(\partial_x B \partial_r G + rB^2(2\partial_r G + r\partial_r^2 G))) + 2r^2 B \partial_r G(\partial_x S - r^2 B \partial_r S)) \\ + S^2(r^4 B^2(\partial_r G)^2 - 2\partial_x G^2) = 0,$$

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Question

How can one describe the dynamics (thermalization) of a (strongly couple) system?

- 1 Analyzing **small excitations** of a uniform static system ...
 - Collective modes in CFT/Quasinormal modes in AdS
 - Janik, Plewa, Spaliński, HS, PRD '15
 - Janik, Jankowski, HS, JHEP '16
 - Janik, Jankowski, HS, PRL '16
 - More in Jakub's talk
- 2 **Full time evolution** of a given initial distribution ...
 - It is interesting but rather complicated!

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Outline

- 1 *AdS/non-CFT setup*
 - How to model the gravity?
 - Numerical setup
 - 1st order phase transition at linear level
- 2 *Full time evolution and 1st phase transition*
 - Gravity setup
 - Time evolution
 - Final state
- 3 *Summary and future directions*

AdS/non-CFT setup

How to model the gravity?

- 1 **Top-down approach:** Deform $\mathcal{N} = 4$ SYM (to $\mathcal{N} = 2^*$) explicitly known (but rather complicated) gravity
- 2 **Bottom-up approach:** Assume AdS/CFT but try to model the gravity+matter background to approach as closely as possible to QCD physics (or other physics of interest)
 - The simplest nontrivial model

$$S = \frac{1}{2\kappa_{d+1}^2} \int d^{d+1}x \sqrt{|g|} \left[R - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - V(\phi) \right] + S_{\text{GH}} + S_{\text{ct}}$$

Gursoy et. al '07; Gubser et. al '08

- We choose $V(\phi)$ such that it reproduces the physics of interest (like lQCD equation of state, 1st or 2nd order transition)

$$V(\phi) \sim -d(d-1) + \frac{m^2}{2} \phi^2 + \mathcal{O}(\phi^4), \quad m^2 = \Delta(\Delta-d)$$

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Numerical set up for static, homogenous solutions

EoS: (Mathematica)

- Einstein equations: **nonlinear** coupled **second order** diff eqs
- Numerical approach: **Newton method**,
spectral **Chebyshev polynomials** with 80 grid points,
numerical precision=**200** and accuracy $\sim 10^{-50}$
- Proper boundary conditions for **asymptotically AdS black hole** solutions in Eddington-Finkelstein coordinates
- The only free parameter is ϕ_H
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$$\epsilon(T) \quad \text{or} \quad s(T) \quad \text{or} \quad c_s^2(T) = \frac{d \ln(T)}{d \ln(s)}$$

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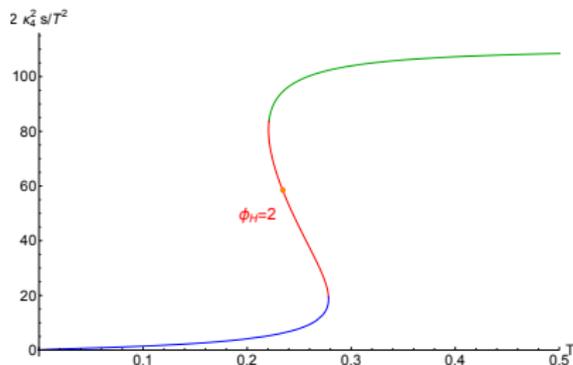
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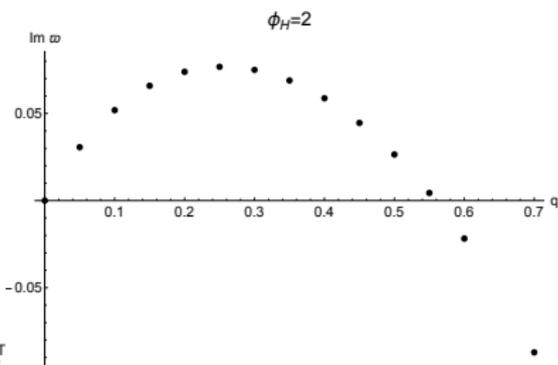
1st order phase transition

Janik, Jankowski, HS '17

EoS



hydro QNM



Thermodynamic instability: $C = T\partial s/\partial T < 0$

Dynamical instability: in large λ for hydro sound/bulk mode

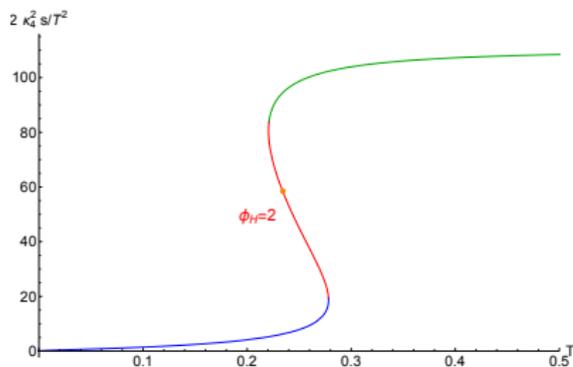
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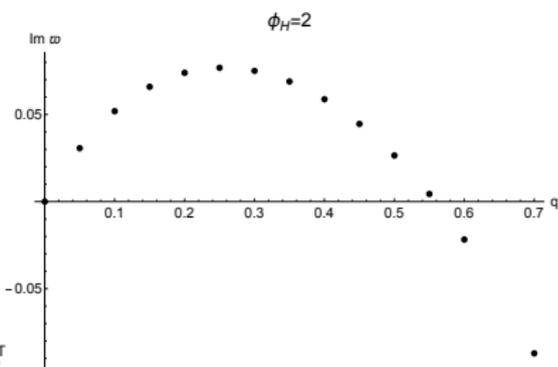
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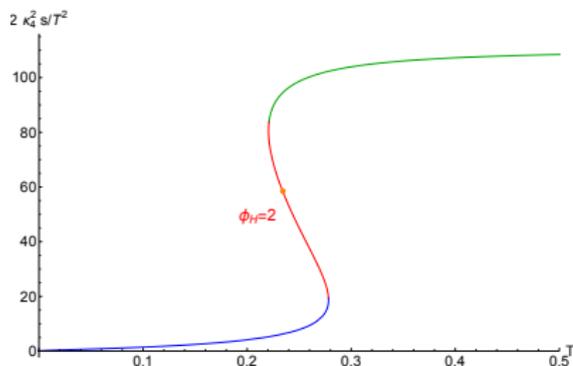
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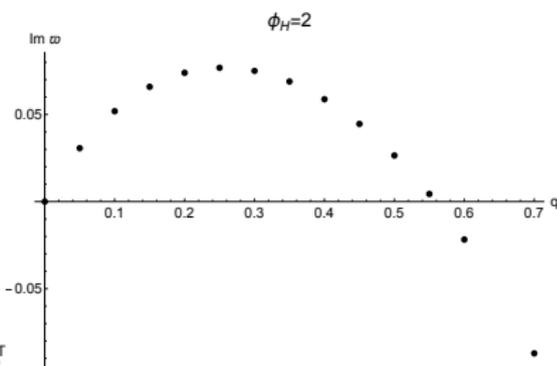
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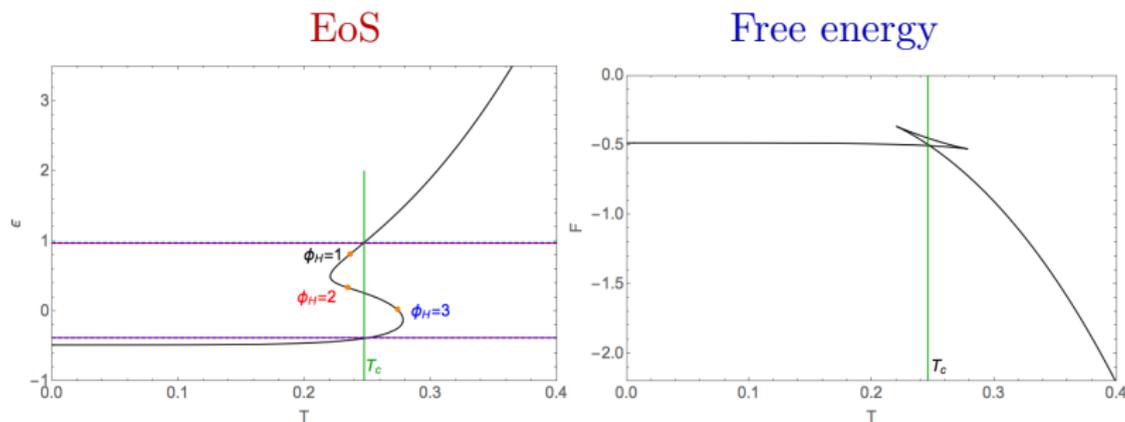
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Nonlinear regime

Dynamical system and 1st phase transition

4D gravity



Potential: $V(\phi) = -6 \cosh\left(\frac{\phi}{\sqrt{3}}\right) - \frac{2}{10}\phi^4$, $\Delta = 2$

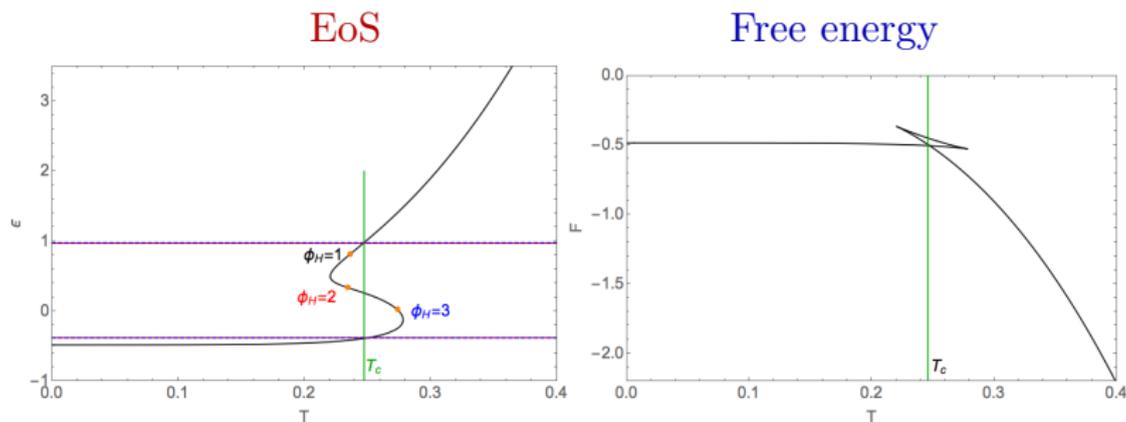
Ansatz (EF coordinate):

$$ds^2 = -A dt^2 - \frac{2 dt dr}{r^2} + S^2 (G dx^2 + G^{-1} dy^2) - 2B dt dx,$$

A, S, G, ϕ are functions of r, t, x .

x is periodic.

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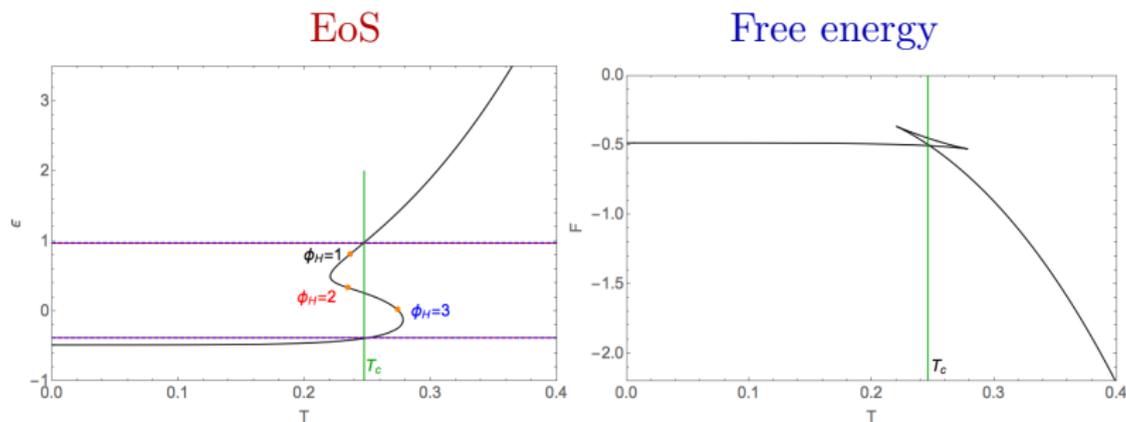
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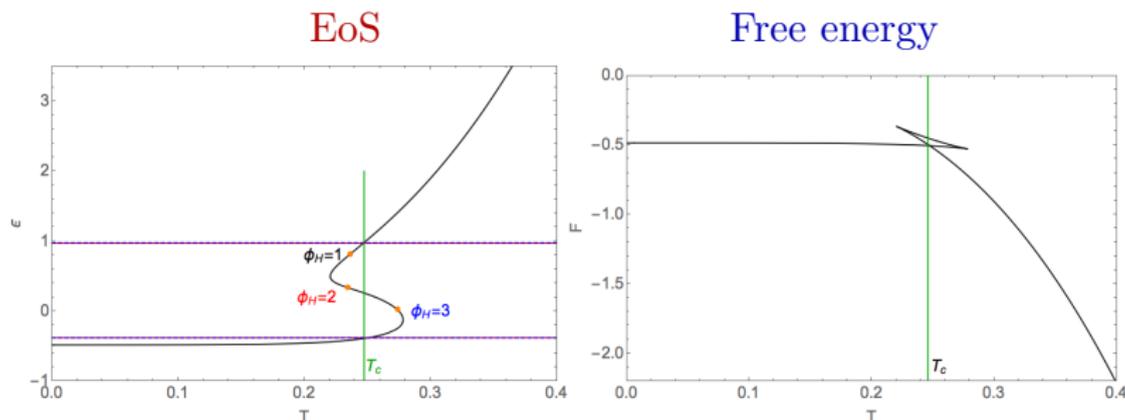
Ansatz (EF coordinate):

$$ds^2 = -A dt^2 - \frac{2 dt dr}{r^2} + S^2 (G dx^2 + G^{-1} dy^2) - 2B dt dx,$$

A, S, G, ϕ are functions of r, t, x .

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Numerical setup for dynamical, inhomogenous solutions

Setup: (Python)

- Einstein equations: **2 + 1** pde's (using $d_+ := \partial_t + \frac{A}{2}\partial_r$)
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solutions with **fixed sources** (Microcanonical ensemble)
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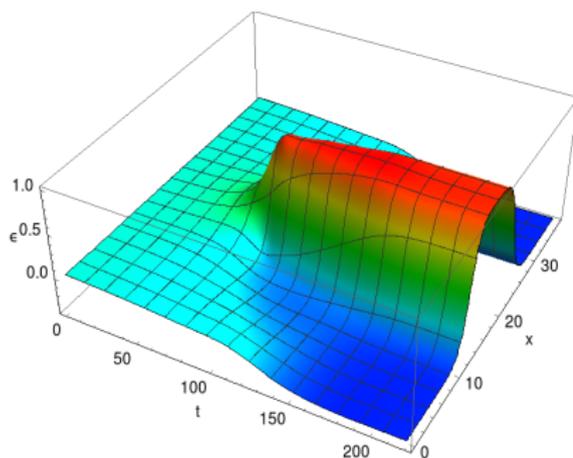
Time evolution

Janik, Jankowski, HS '17

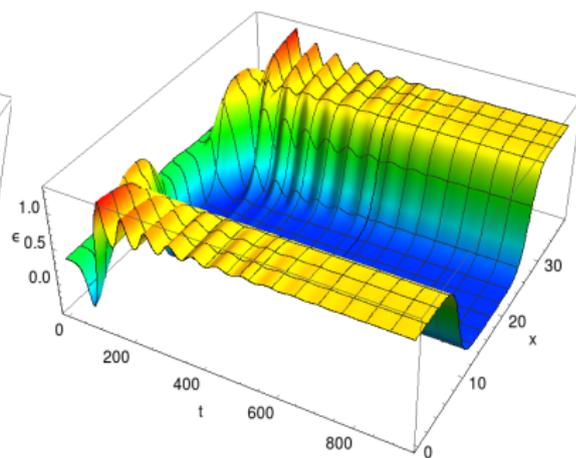
Perturbations with 12π period: Periodic solution

$$\delta S \propto \cos(kx)$$

$$\delta S \propto \exp(-w_0 \cos(kx)^2)$$



$$\phi_H = 3$$



$$\phi_H = 2$$

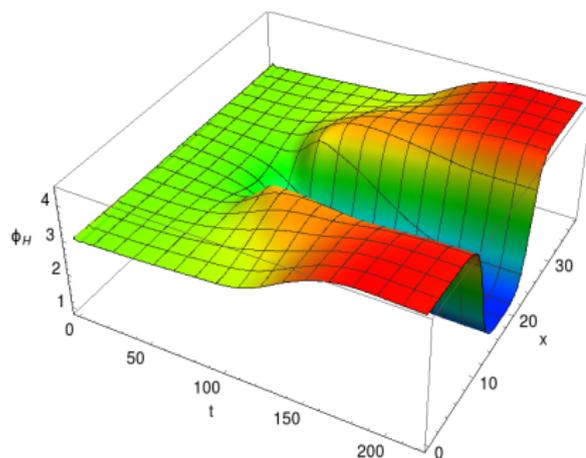
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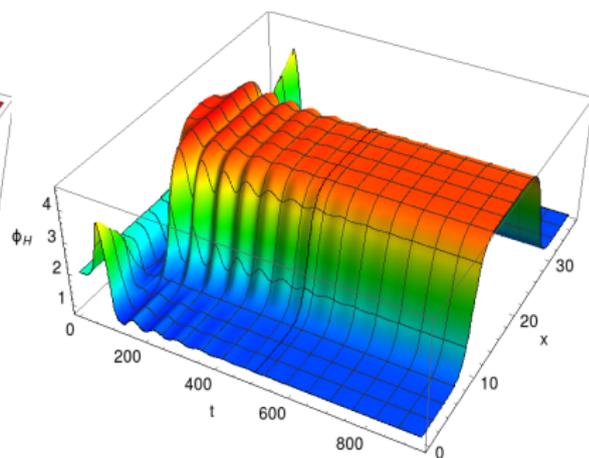
Perturbations with 12π period: Inhomogenous horizon

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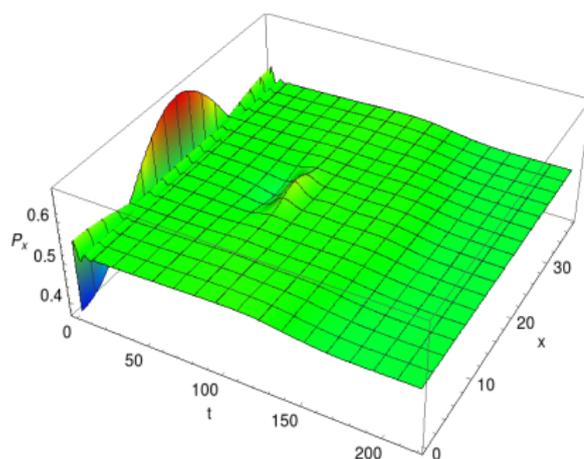
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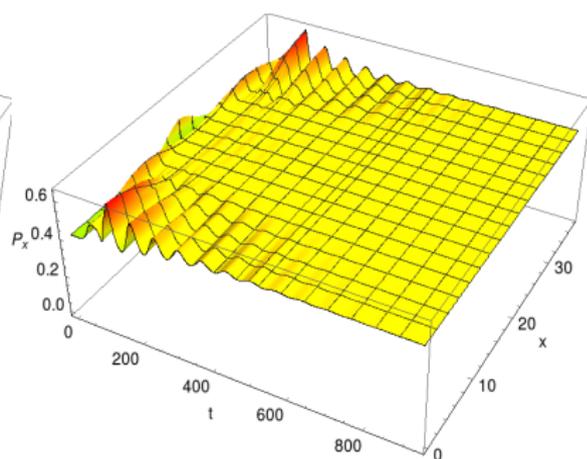
Perturbations with 12π period: No gradient along x at late time

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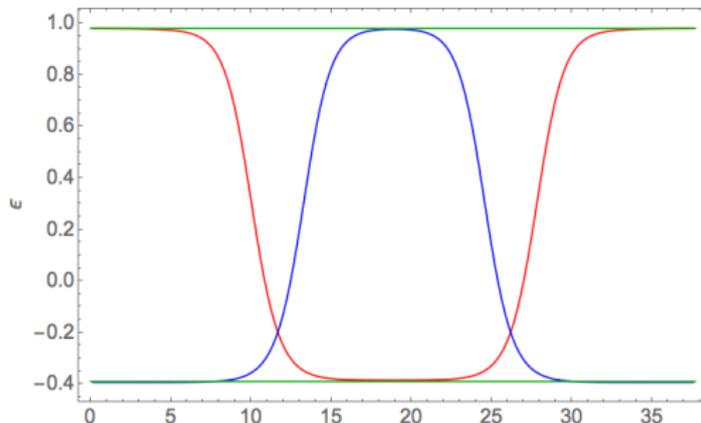


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Final state

Janik, Jankowski, HS '17

Final state seats at $T = T_c$:



Green lines: Min and Max points on EoS at T_c

Blue line: $\epsilon(x, t_f)$ for the single mode perturbation

Red line: $\epsilon(x, t_f)$ for mixed modes perturbation

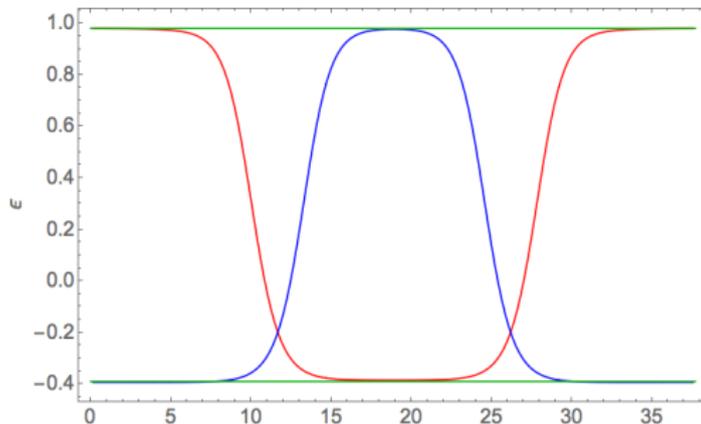
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$$P_{x/y} = P_{eq}(\epsilon) + c_{x/y}(\epsilon) (\partial_x \epsilon)^2 + f_{x/y}(\epsilon) (\partial_x^2 \epsilon)$$

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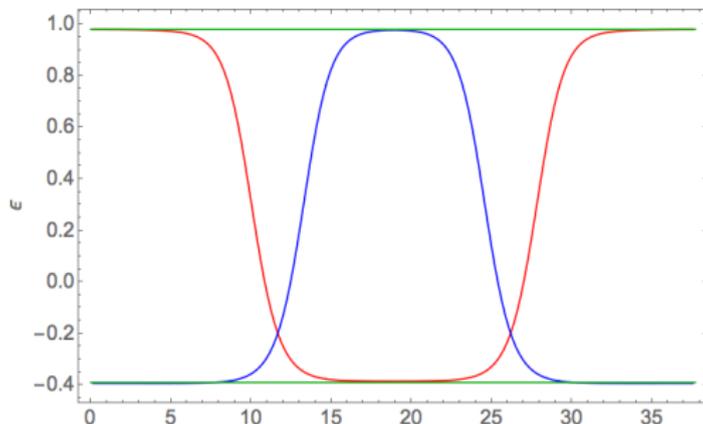
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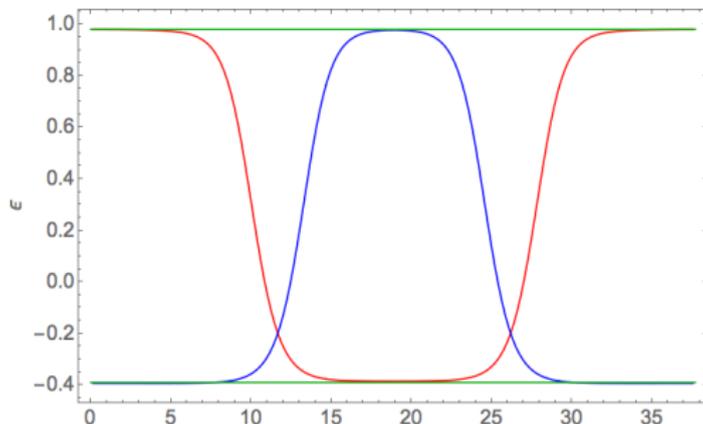
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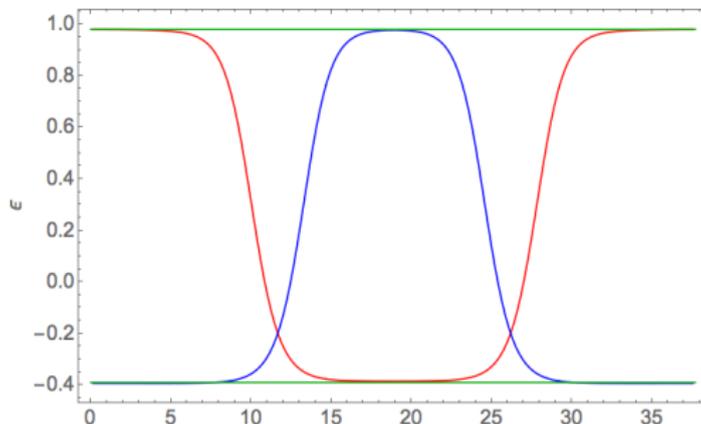
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- 2 There are new stationary **domain wall** BHs
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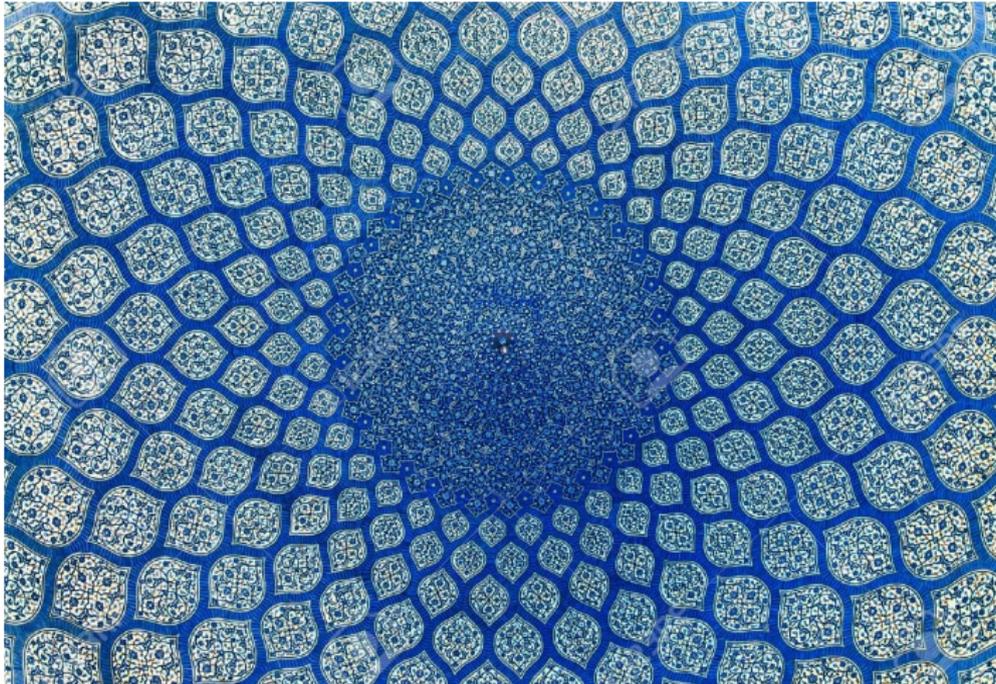
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Thank you for your attention

Backups

