



Strolling along gravitational vacua

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- Strominger's infrared triangle



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- 1. Warm-up: Particle theory
- 2. Maxwell theory and its vacua
- 3. Gravitational vacua and low energy dynamics

Warm-up: Particle theory

$$L = \frac{1}{2}g_{IJ}(X)\dot{X}^I\dot{X}^J - V(X^I)$$

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- \blacktriangleright Space of vacua $\mathcal{V}=S^2=SO(3)/SO(2)$
- Spherical coordinates $(R + \rho, z^{\alpha})$

$$(R+\rho)\left(\dot{z}^{\alpha}+\Gamma^{\alpha}_{\beta\gamma}\dot{z}^{\beta}\dot{z}^{\gamma}\right)+2\dot{\rho}\dot{z}^{\alpha}=0$$
$$\ddot{\rho}+\frac{4}{m}\rho(\rho+2R)(\rho+R)-(R+\rho)g_{\alpha\beta}\dot{z}^{\alpha}\dot{z}^{\beta}=0$$

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► Low energy solutions: Geodesics on the space of vacua.

Maxwell theory and its vacua

• Temporal gauge $A_0 = 0$

$$L[A_i] = \frac{1}{2} \int d^3x \left(g^{ij} \dot{A}_i \dot{A}_j - F_{ij} F^{ij} \right)$$

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 - Metric on the configuration space

$$g(\delta_1 A, \delta_2 A) = \int d^3 x \operatorname{Tr} \delta_1 A_i \, \delta_2 A^i$$



Vacuum configurations

 $\mathcal{V} = \min(V)$

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- Allowed directions (LGTs)

$$\phi_{\alpha} = (\frac{r}{R})^{\ell} Y_{\ell m} \qquad \mathsf{H}$$

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Theorem:

$$\mathcal{G}_{\mathsf{harmonic}} \cong \mathcal{G}_{\mathsf{boundary}} \cong \mathcal{G}/\mathcal{G}_0$$

• Geodesic equation $\ddot{z}^{\ell m}(t) = 0$

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- Metric on the space of vacua

$$g(\nabla \phi_{\ell m}, \nabla \phi_{\ell' m'}) = \int_{M} \nabla_{i} \phi_{\ell m} \nabla^{i} \phi_{\ell' m'}$$
$$= \oint_{\partial M} \phi_{\ell' m'} n \cdot \nabla \phi_{\ell m} = \ell \, \delta_{\ell \ell'} \, \delta_{m m'}$$

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Physical vs. pure gauge

$$\mathbf{g}(\nabla \phi_{\mathsf{harm}}, \nabla \lambda_{\mathsf{pure gauge}}) = 0, \qquad \lambda_{\mathsf{pure gauge}}\Big|_{\partial M} = 0$$

Adiabatic approximation as a probe of vacua and symmetries

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- ► This is missing in asymptotic symmetry analyses

Gravitational vacua and low energy dynamics

• Einstein gravity
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Field equations

$$\begin{split} R_{ij} + \frac{1}{2}\ddot{h}_{ij} - \frac{1}{2}h^{kl}\dot{h}_{k[i}\dot{h}_{l]j} &= 0 \qquad \text{Dynamical equation} \\ \nabla^{i}\left(\dot{h}_{ij} - h^{kl}\dot{h}_{kl}h_{ij}\right) &= 0 \qquad \text{Momentum constraint} \\ R + \frac{1}{2}h^{ij}h^{kl}\dot{h}_{i[j}\dot{h}_{k]l} &= 0 \qquad \text{Hamiltonian constraint} \end{split}$$

Configuration space

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• Coordinate system z^{α} $\bar{h}(x;z) = g_z \cdot h_{\text{ref}}$

Manton approximation

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Velocity (extrinsic curvature)

$$\dot{h} = \delta_{\chi_z} \bar{h} = \bar{\nabla}_{\rm s} \chi_z \qquad \chi_z^i = \dot{\phi}_z^k \phi_z^i{}_{\underline{k}}$$

Low energy dynamics

$$S[z(t)] = \int dt \frac{1}{2} \bar{\mathbf{g}}_z(\dot{z}, \dot{z}) \qquad \bar{\mathbf{g}}_z(\dot{z}, \dot{z}) = \mathbf{g}(\bar{\nabla}_{\mathbf{s}} \chi_z, \bar{\nabla}_{\mathbf{s}} \chi_z)$$

Momentum constraint

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Metric on the space of vacua

$$\begin{aligned} \langle \chi_1, \chi_2 \rangle_h &\equiv g(\nabla_s \chi^{(1)}, \nabla_s \chi^{(2)}) \\ &= \oint_{\partial M} \sqrt{k} \, d^2 y \left(\chi^a_{(1)} D^\perp \chi^{(2)}_a - K_{ab} \chi^a_{(1)} \chi^b_{(2)} \right) \end{aligned}$$

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• Hamiltonian constraint $\langle \chi_1, \chi_2 \rangle_h = 0$

- Bulk extension of $\chi \Big|_{\partial M} = \zeta$ is unique up to exact vector fields $\partial_i \phi$.
- ▶ This however, does not affect the metric
- Conjecture The Hamiltonian constraint then fix the remaining part uniquely
- All the data is available on the boundary

Conserved momenta

Conserved momenta

$$P_{\zeta} \equiv g_z(\dot{z}, \delta_{\zeta} h) = \langle \zeta, \chi_z \rangle_{\bar{h}(z)} = \int_{\partial M} \sqrt{\bar{k}} d^2 y \, n^i e_a^j \zeta^a \dot{h}_{ij}$$
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Surface charges from covariant phase space

$$Q_{\chi} = \int_{M} \Theta(\delta_{\chi} h)$$

$$\Theta(\delta_{\chi}h) = \int_{M} d^{3}x \sqrt{h}(\dot{h}_{ij} - h_{ij}h^{kl}\dot{h}_{kl})\nabla^{i}\chi^{j}$$
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► The momenta are equal to Noether surface charges

$$P_{\zeta} = Q_{\chi}$$
 where $\zeta^a = \chi^a|_{\partial M}$

Diffeos on the boundary

$$\zeta_a = \sqrt{k} \epsilon_a{}^b \partial_b \tau + \partial_a \rho \qquad \tau = \sum_{\ell,m} a_{\ell m}, \quad \rho = \sum_{\ell,m} b_{\ell m} Y_{\ell m}$$

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- ► Bulk extension $\chi_i = \eta_i + \partial_i \phi, \qquad \chi_a|_{\partial M} = \zeta_a$
- Momentum constraint

$$\eta = r \times \nabla H, \qquad H = \sum_{\ell,m} a_{\ell m} \left(\frac{R}{r}\right)^{\ell+1} Y_{\ell m}$$

Diffeos on the boundary

$$\zeta_a = \sqrt{k} \epsilon_a{}^b \partial_b \tau + \partial_a \rho \qquad \tau = \sum_{\ell,m} a_{\ell m}, \quad \rho = \sum_{\ell,m} b_{\ell m} Y_{\ell m}$$

- ► Bulk extension $\chi_i = \eta_i + \partial_i \phi, \qquad \chi_a|_{\partial M} = \zeta_a$
- Momentum constraint

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Metric on the space of vacua

$$\langle \zeta_{(1)}, \zeta_{(2)} \rangle = R^3 \sum_{\ell,m} \ell(\ell+1) \left((\ell-1) a_{\ell m}^{(1)} a_{\ell m}^{(2)} - 2 b_{\ell m}^{(1)} b_{\ell m}^{(2)} \right)$$

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► Hamiltonian constrain

$$\sum_{\ell,m} (\ell - 1)a_{\ell m}^2 - 2b_{\ell m}^2 = 0$$

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Modeling a generic stationary solution (conjecture)

- ► Adiabatic approximation as a probe of the space of vacua
- ▶ It allows to identify the physical symmetries and their bulk extensions
- and to identify a set of solutions of the theory
- Divergences in the large volume limit
- Solve the bulk extension and identify solutions

Thank you for your attention